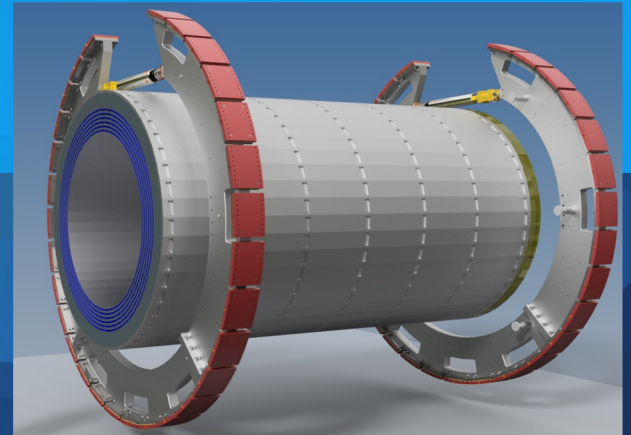


# Barrel Imaging Calorimeter for ePIC at the EIC

Sanghoon Lim  
Yonsei University



# HERA vs EIC: key parameters

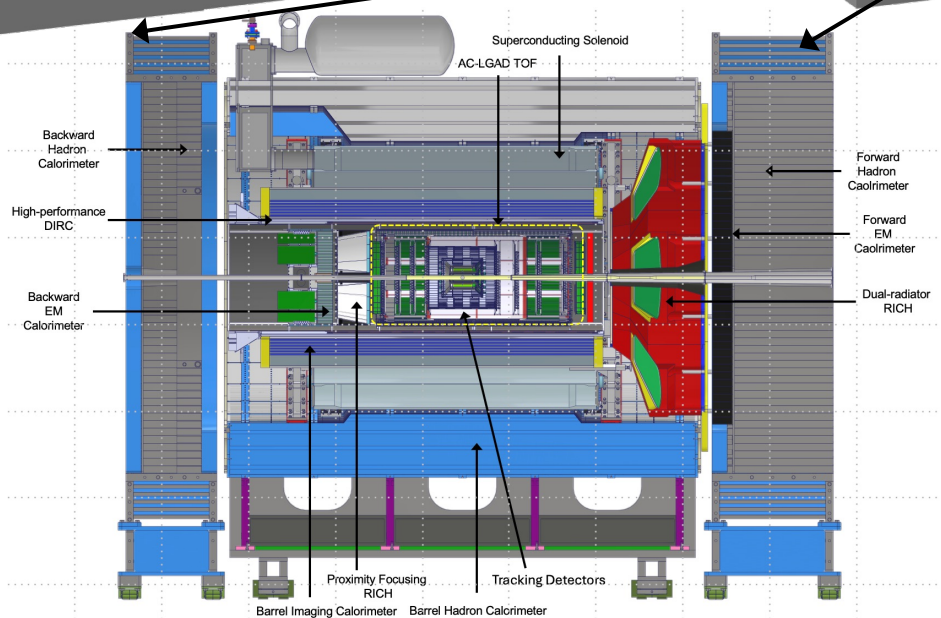
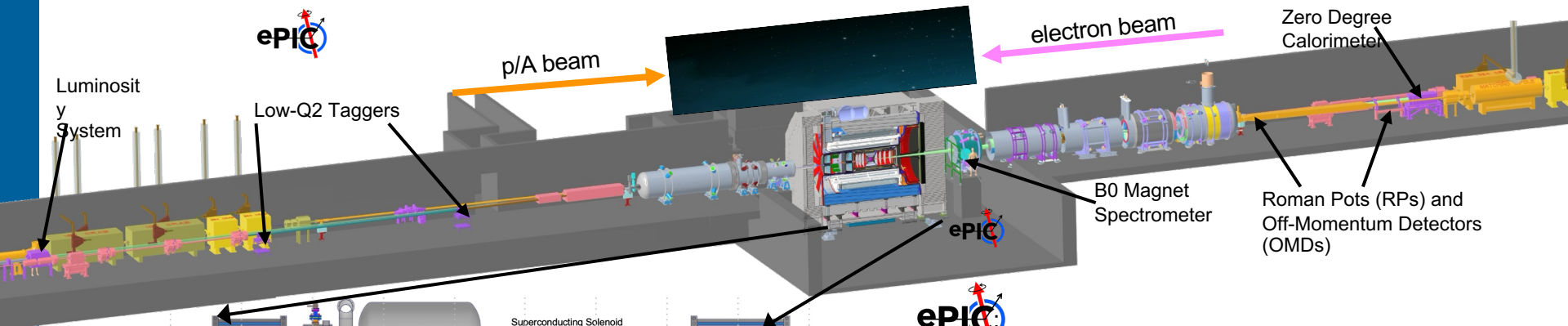
	HERA	EIC	FACTOR
Circumference (km)	6.3	3.8	1.7
Number of bunches	174	1160	6.7
Proton bunch charge (nC)	11.7	11	1
Electron bunch charge (nC)	5.3	28	5.3
Bunch length (cm), p/e	16/0.9	6/0.7	
Beta at IP, proton H/V (cm)	245/18	80/7.2	
Beta at IP, electron H/V (cm)	62/26	45/5.6	
Proton emittance, (nm, rms)	4/4 ( <i>round</i> )	11/1 ( <i>flat</i> )	
Electron emittance, (nm, rms)	20/3 ( <i>flat</i> )	20/1.3 ( <i>flat</i> )	
Energy (CM), GeV	320	105	
<b>Peak Luminosity, <math>\times 10^{31} \text{ cm}^{-2}\text{s}^{-1}</math></b>	<b>5.3</b>	<b>1000</b>	<b>190</b>

} 60  
} ~3

$$L = \frac{N_p N_e f_0}{4\pi\sigma_x\sigma_y} \approx 1 \times 10^{31} \text{ cm}^{-2}\text{s}^{-1}$$

Electron-Ion Collider

# ePIC – The EIC General-Purpose Detector



26 subsystems over  $\pm 40$  m

to measure particle momenta, energy and particle type

- 3 electromagnetic calorimeters
- 3 hadronic calorimeters
- Silicon and Multi-pattern gas detectors
- 3 RICH detector + time-of-flight
- 7 Auxillary detectors (Si + HCal + ECals)
- electron and hadron polarimetry

Integration, Installation and Infrastructure

Non-Beam Commissioning

Highest scientific flexibility

- fully streaming readout electronics and data acquisition
- integration AI/ML capabilities from the start

# Detector Requirements for BIC

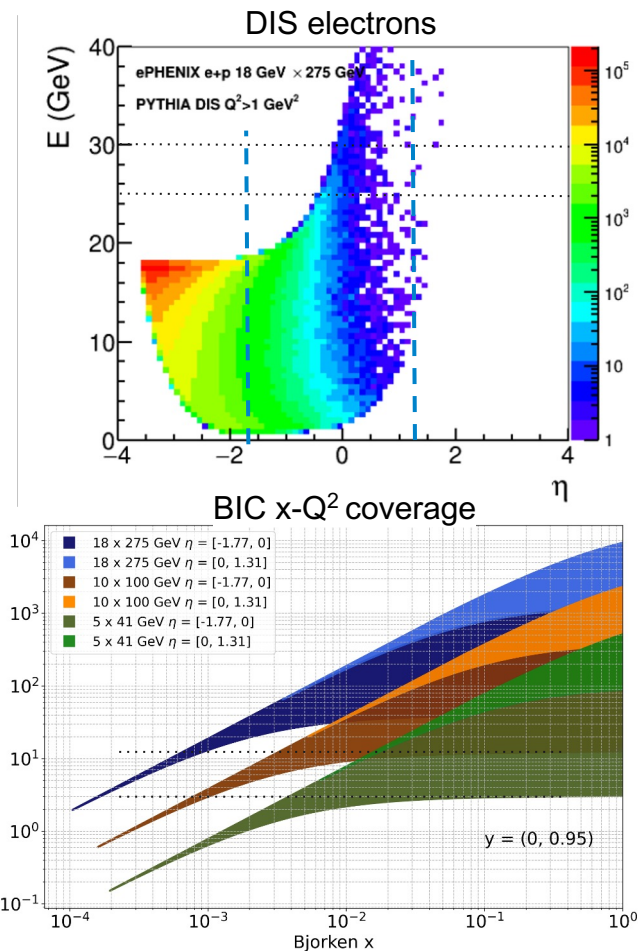
<https://eic.jlab.org/Requirements/>

Identify scattered electrons and measure their energy, in high  $Q^2$  events, also decay electrons, e.g., from vector or heavy flavor meson decays, and measure DVCS photons and decay photons

- **Electron ID up to 50 GeV** and down to 1 GeV and below
  - Energy resolution  $< 10\%/\sqrt{E} + (2-3)\%$
  - High power for  **$e/\pi$  separation down to 1 GeV/c**
- **Photon measurements up to 10 GeV**
- **$\gamma/\pi^0$  separation up to 10 GeV**
  - Distinguishing two showers with an opening angle down to 30 mrad

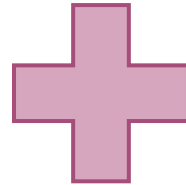
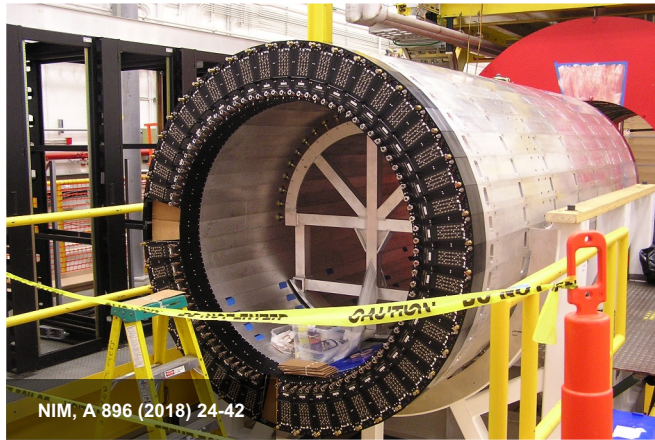
**Assist with muon identification**

Sufficient dynamic range to **detect MIP** signals in all layers



# A Hybrid Imaging Calorimeter

Combination of a high-performance sampling calorimeter with silicon sensors for shower profiling



Start from mature layered Pb/ScFi technology with side-readout (same as the GlueX calorimeter) for state-of-the-art sampling calorimeter performance

Insert layers of monolithic AstroPix sensors (ultra-low-power silicon sensor developed for NASA) in the first half of the calorimeter to capture a 3D image of the developing shower

# Pb/ScFi Layer Technology

## Our Pb/ScFi layers follow the GlueX Design

Energy resolution at GlueX:  $\sigma = 5.2\% / \sqrt{E} \oplus 3.6\%$ <sup>1)</sup>

- GlueX has  $15.5 X_0$ , and could not constrain the constant term (due to low energies) → fixed in FTBF/CERN Beam Test (< 2%)

Position resolution in z:  $1.1 \text{cm} / \sqrt{E}$ <sup>2)</sup>

- 2-side SiPM readout,  $\Delta t$  measurement

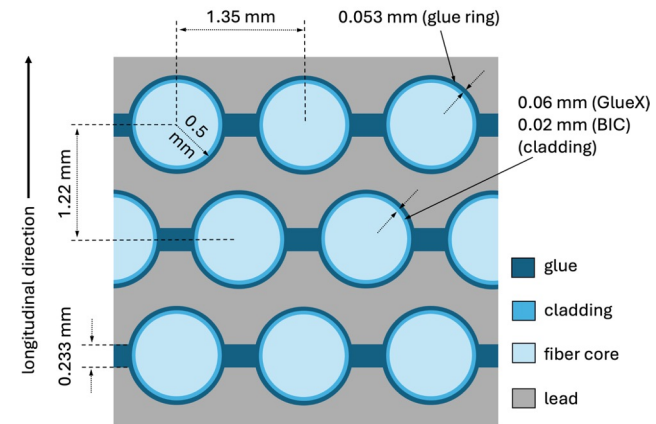
**Mature technology** used in Barrel ECals (GlueX, KLOE)

- Detailed studies on **calorimetry performance**, including the light collection uniformity in fibers, light collection efficiencies, etc.
- **Construction** (lead handling, swaging, Pb/ScFi layers assembly, module machining) is fully developed for GlueX

Z. Papandreou, <https://haldweb.jlab.org/DocDB/0031/003164/>

**Long-lead procurement items are ongoing for BIC (past FDR)**

- SiPMs  $1.2 \times 1.2 \text{ mm}^2$  arrays, 50  $\mu\text{m}$  pitch (Hamamatsu S14160-6050-04 or similar)
- Single Clad Kuraray (SCSF-78,  $\phi$  1mm)



- 1) Nucl. Instrum. Meth. A, vol. 896, pp. 24–42, 2018
- 2) Nucl. Instrum. Meth. A, vol. 596, pp. 327–337, 2008

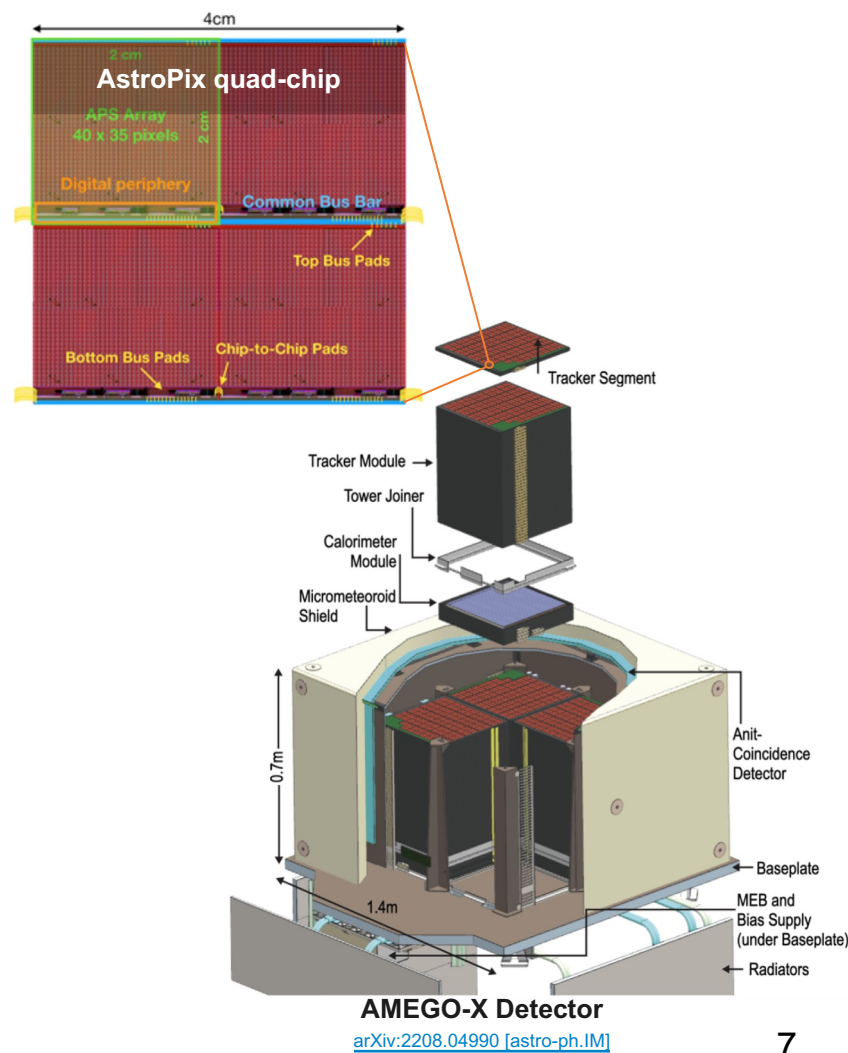
# Imaging Layer Technology

## Imaging layers will use the **AstroPix sensors**

- Developed for NASA AMEGO-X space mission  
[arXiv:2109.13409](https://arxiv.org/abs/2109.13409) [astro-ph.IM]

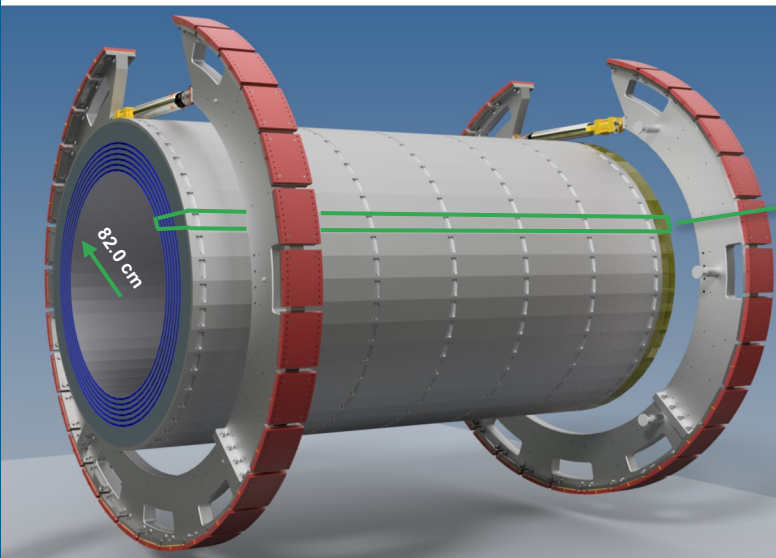
## Key features:

- Very low power dissipation - will be used in space!
- Good energy resolution (thick silicon sensor)
- 500  $\mu\text{m}$  pixel size ( $\sim 144 \mu\text{m}$  resolution)
- **Perfect for use in calorimetry!**
- First silicon layer has sufficient resolution to be used as a tracking layer behind the DIRC (replacing the MPGD layer)



# Barrel Imaging Calorimeter

## Sector Mechanics

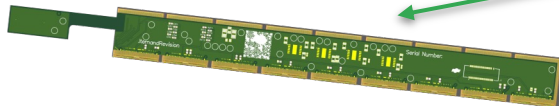


Total BIC weight ~42.5 US tons

**AstroPix Module** - Nine AstroPix sensors daisy-chained together on Flex PCB.

A stave consists of 12 modules.

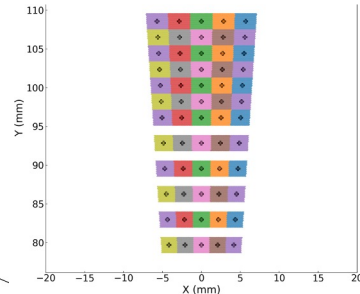
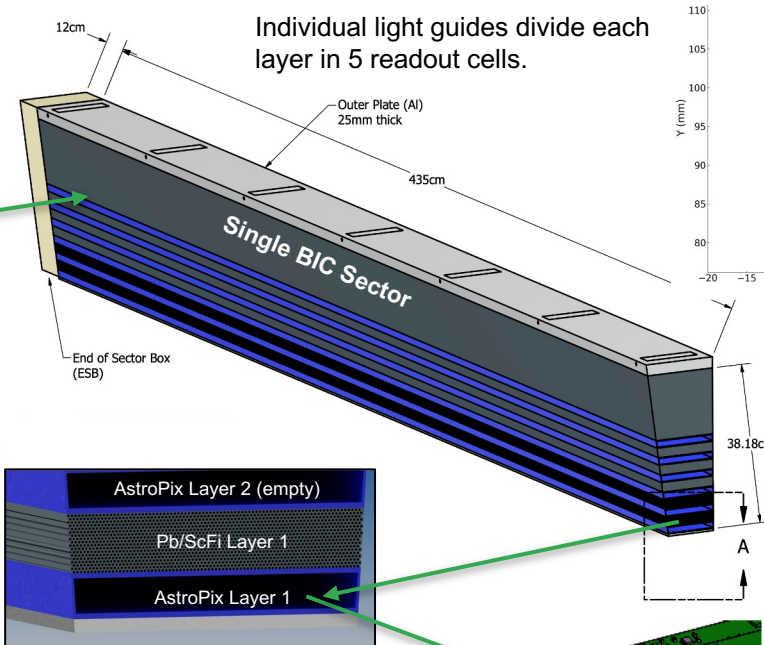
A tray contains of 6-8 staves.



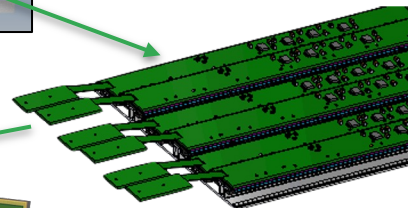
**Pb/SciFi Layers** - 17 rows of fiber between corrugated lead.

Each sector has 12 Pb/SciFi layers.

Individual light guides divide each layer in 5 readout cells.



**SiPMs** - 5 per layer  
x 12 layers = 60  
(Hamamatsu S14161-3050)

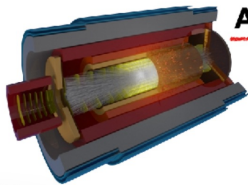
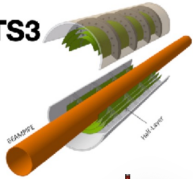


**Tray** - Structure holding the AstroPix staves for a single layer (217.5 cm long).

# Synergy with other detector projects

## Silicon Tracker in ALICE

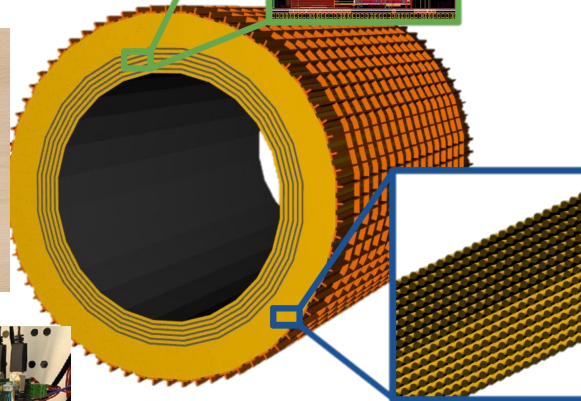
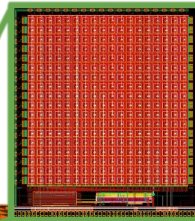
ITS3



ALICE 3

**Position resolution:**  
Silicon layers  
with AstroPix  
 $500 \times 500 \mu\text{m}^2$  pixel size

## ePIC BIC



**Energy resolution:**  
Pb/SciFi layers  
with two-sided  
SiPM readout

## Dual-Readout Calorimeter

### Innovative Detector for e+e- Accelerator (IDEA) detector concept with DRC

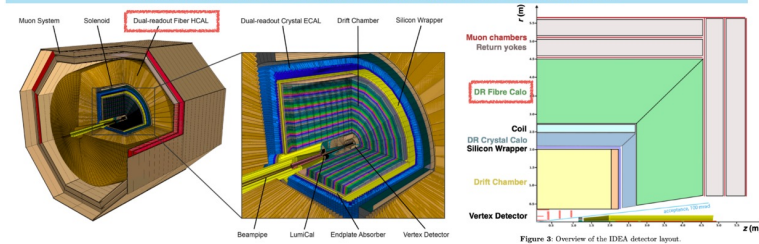
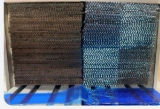


Figure 3: Overview of the IDEA detector layout.


**TB2022**



Cu Plate

**EM Performance Study**  
Cu Plate  
 $-3.5 \rho_{\text{tr}} > 30 X_0$   
PMT, SiPM  
 $10 - 120 \text{ GeV } e^+e^-$   
@SPS H8

**TB2023**



3D-Printed Module, SFHS, Lego-Like

**Low Energy EM Performance Study**  
3D-Printed Module, SFHS, Lego-Like  
 $-3 \rho_{\text{tr}} \sim 17 X_0$   
PMT, MCP-PMT  
 $0.5 - 5 \text{ GeV } e^+e^-$   
@PS T9

**TB2024**

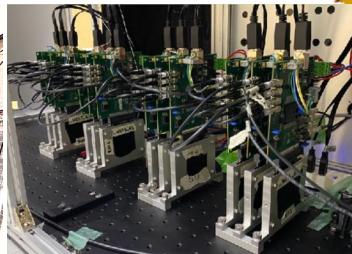


Full Size Hadron Deposit

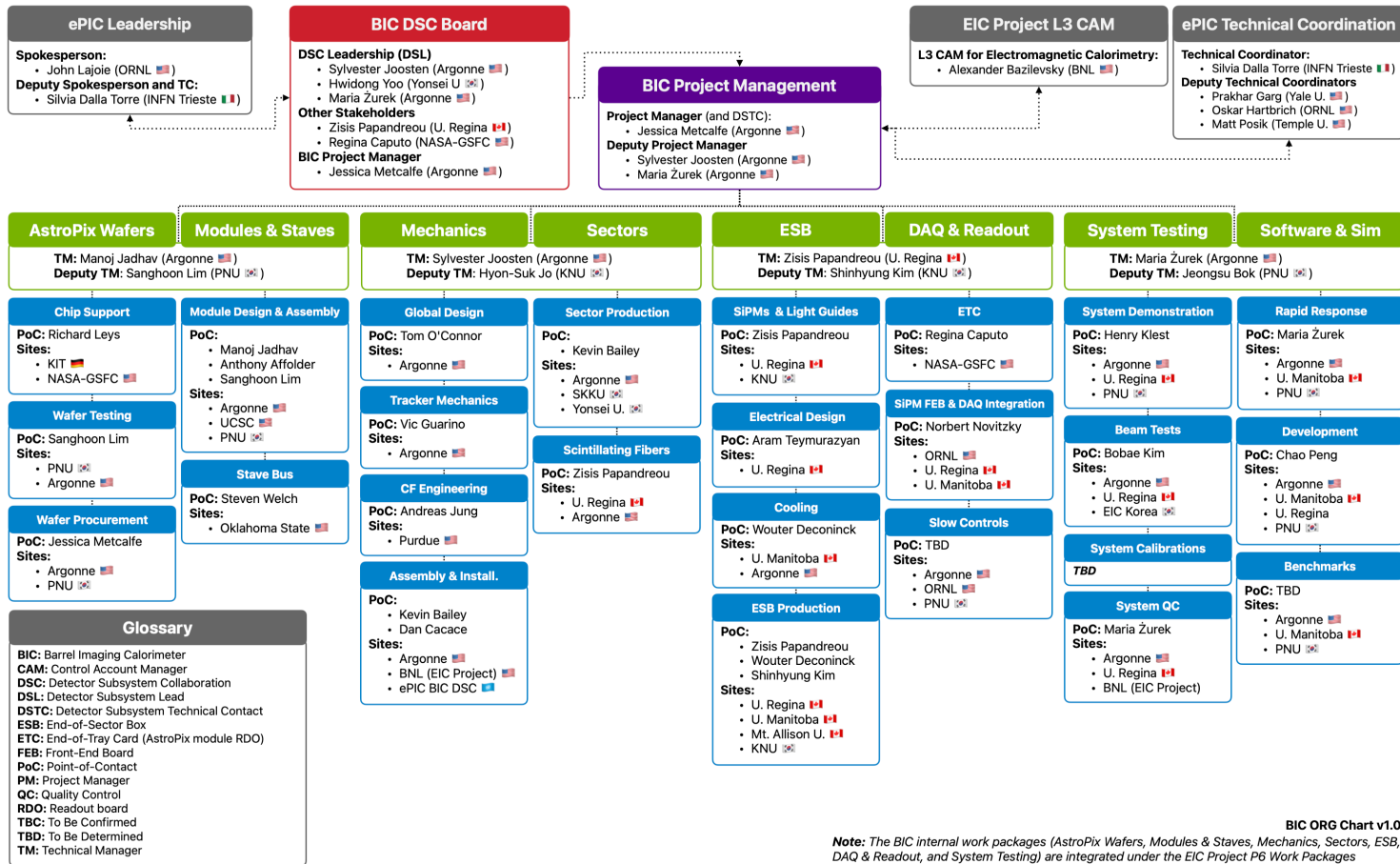


Skinned Fish Head Sil

**EM&Hadron Performance Study**  
SFHS, Cu Plate  
 $\sim 90\%$  lateral hadron deposit,  $\sim 10 \mu\text{m}$   
PMT, MCP-PMT, SiPM  
 $10 - 120 \text{ GeV } e^+e^-$ ,  $20 - 120 \text{ GeV } \mu^+\mu^-$   
@SPS H8



# BIC Organization Chart



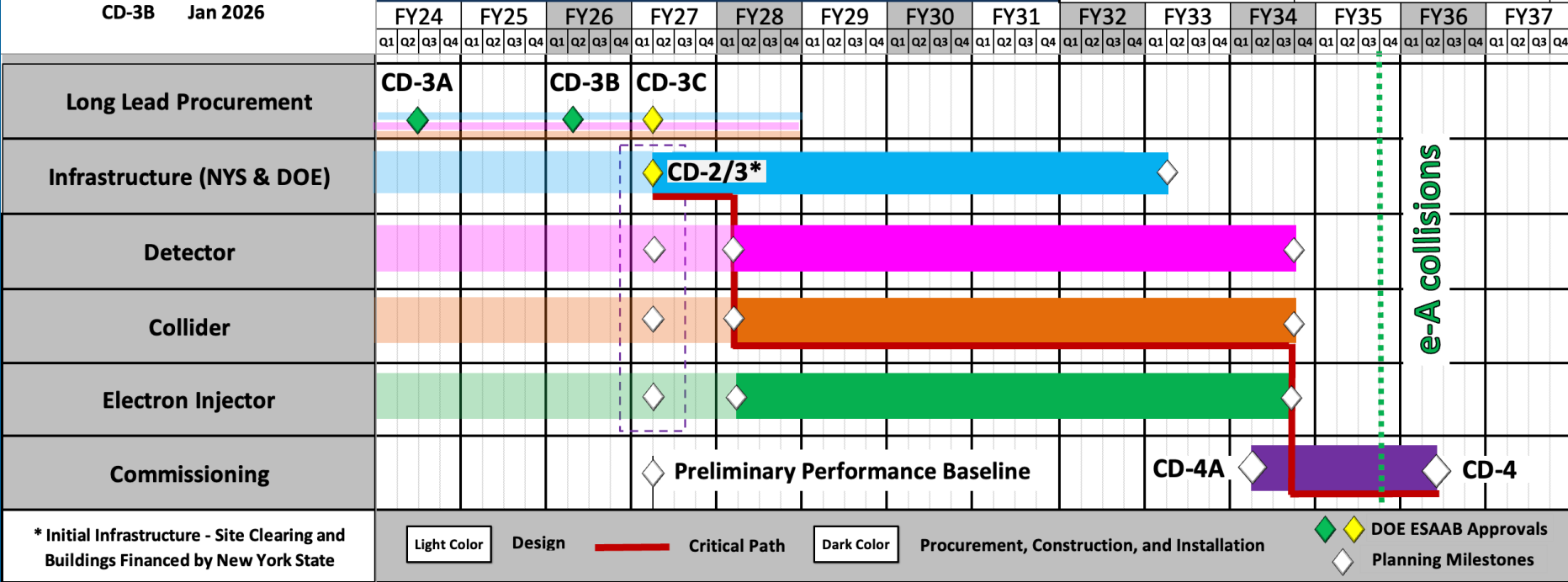
# EIC Project Schedule

DOE ESAAB Approvals

- CD-1 Jun 2021
- CD-3A Mar 2024
- CD-3B Jan 2026

## Electron-Ion Collider

Fiscal Year (FY) = Oct 1 - Sep 30

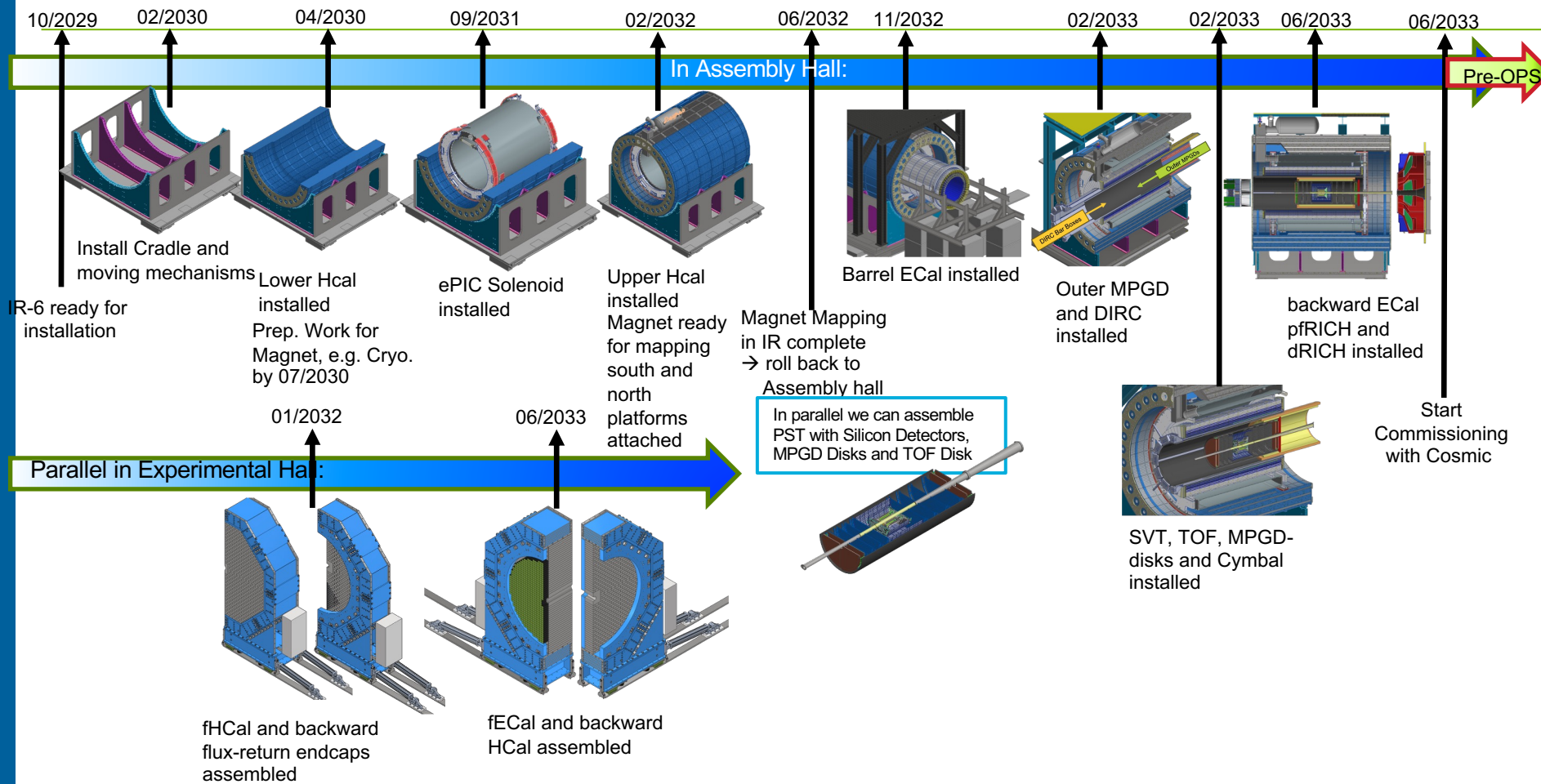


e-A collisions

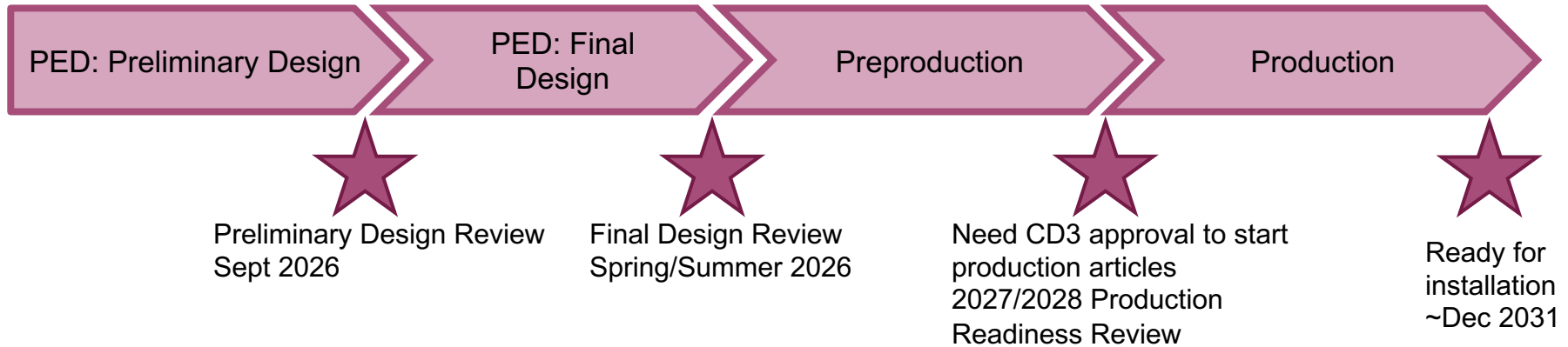
Light Color Design    
  Critical Path    
 Dark Color Procurement, Construction, and Installation    
 ◆ ◆ DOE ESAAB Approvals    
 ◆ Planning Milestones

From the 7th EIC RRB meeting on June 9-10

# Detector Installation Schedule



# BIC Project Phases (to be updated)



- Preliminary Design Review, 60% design completion
- AstroPix v3 (and v4)
- BabyBCal & Lanky BCal
- Individual components
- First (second) test articles

- AstroPix v5
- One full sector
- Final designs (90%)
- Production style procedures
- AstroPix v6 validation tests

- AstroPix v6
- 48 sectors

# Production Plan

## Parts

	Total on-detector	Yield at step	Cumulative Yield	Total # parts to produce
sectors	48			
trays	384	99%	99%	388
staves	2592	98%	97%	2,672
modules	31104	95%	92%	33,747
AstroPix chips	279936	85%	78%	357,319
wafers				5,956
end-of-tray cards	384	90%	90%	427
tray support rails	384	95%	95%	405
module PCBs	31104	90%	81%	41,663
module base plate	31104	95%	87%	35,523
fiber canes	1M	98%	98%	4900 km
end-of-sector-boxes	96	96%	96%	100
SiPM's	5760	96%	96%	6022
CALOROC PCB	96	95%	95%	101

~100 m<sup>2</sup> of silicon

~5,000 km of fibers

Yields are based on experience with existing silicon detector productions and GlueX

will need to be verified to improve estimate maturity

# Production Plan

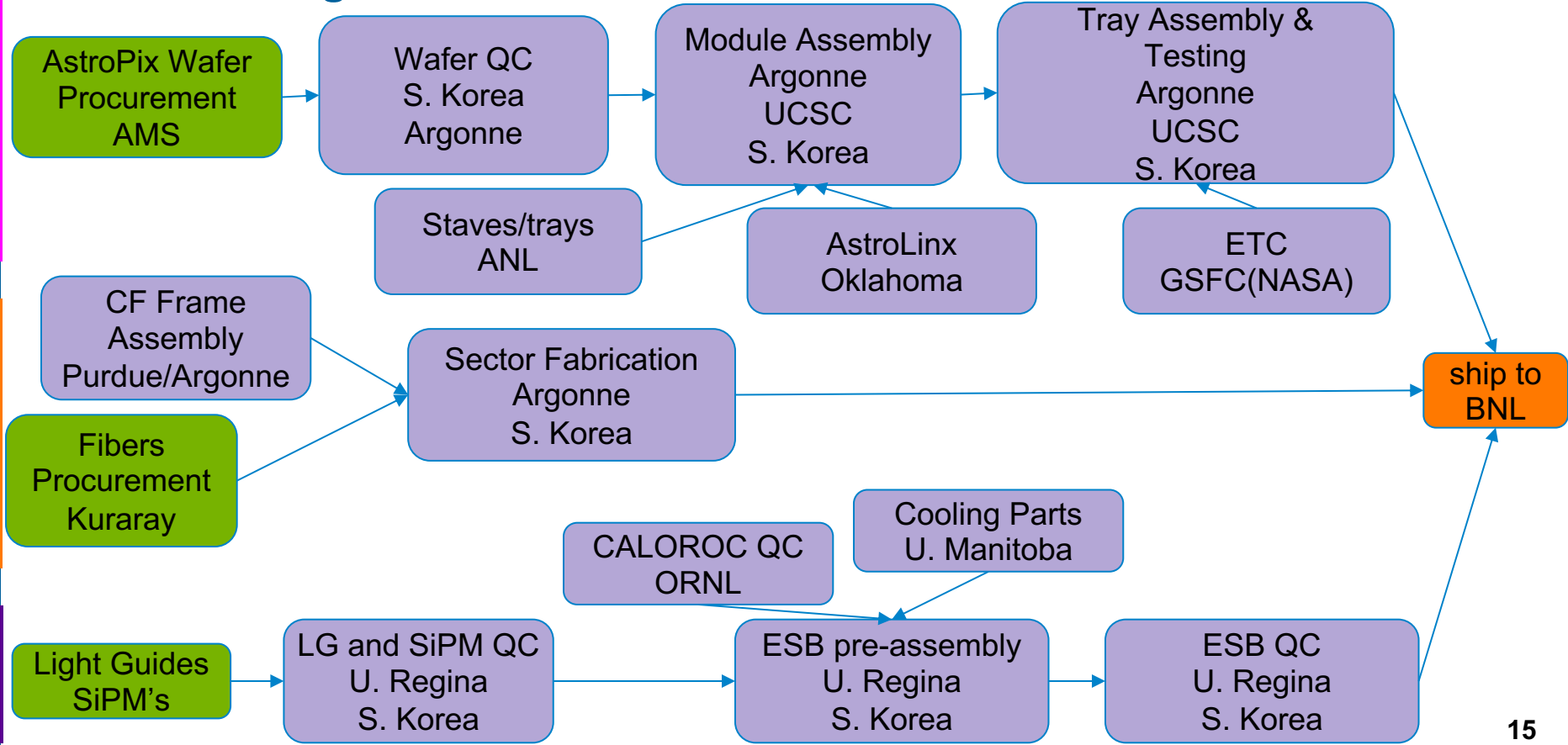
## Plan including Korean in-kind contribution\*

vendor  
site

Tracker

Sector

ESB



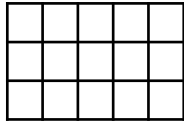
# Activities in PED



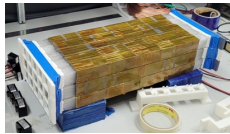
May 2024



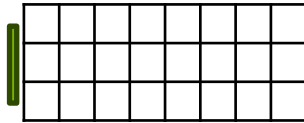
First piece



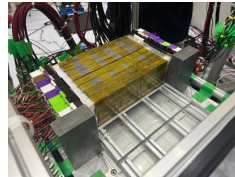
Aug 2024 CERN PS



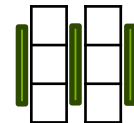
First Pb/SciFi prototype  
production in Korea  
First beam test  
**Successful data taking**



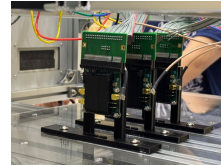
Mar 2025 KEK



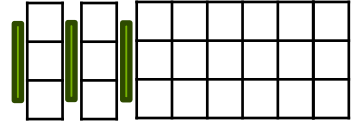
More Pb/SciFi layers  
Improved energy resolution  
**New DAQ, Trigger,**  
Hodoscope, Calibration  
**Extension cable test**



June 2025 KEK



**AstroPix synchronization**  
with trigger time

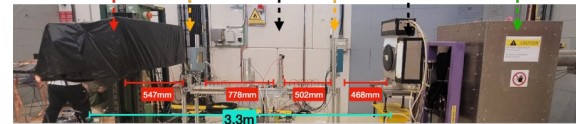
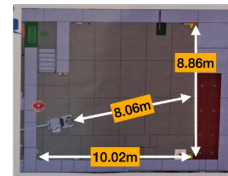
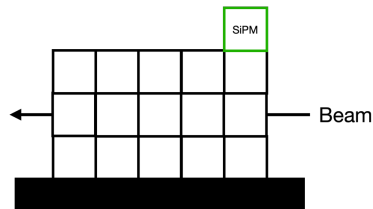
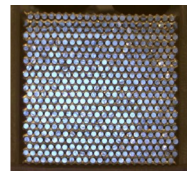
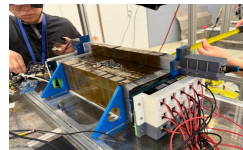
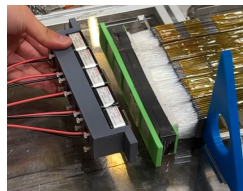
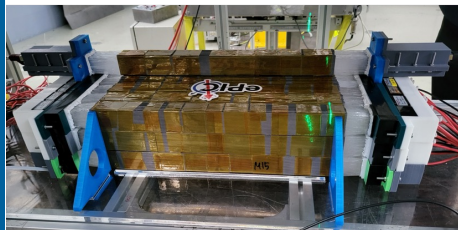
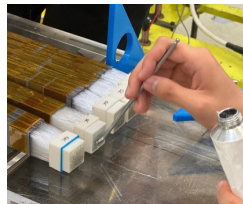
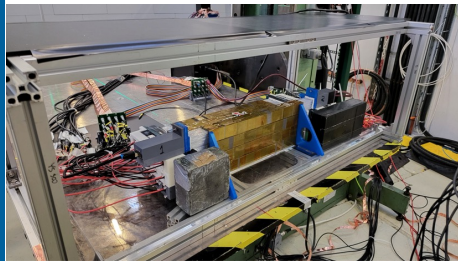


July 2025



**synchronized data taking of  
AstroPix + Pb/SciFi  
Proof-of-principle imaging of  
electrons and pions**  
Resolution fulfill the EIC  
requirement

# First beam test at CERN PS T10 in Aug 2024



Successfully collected data from the Pb/SciFi calorimeter prototype produced in Korea using our DAQ system with 0.5, 1, 2, and 3 GeV electron beams.

arXiv:2604.22647

Beam test of a Pb/SciFi prototype for the Barrel Imaging Calorimeter at the Electron-Ion Collider

Hyungjun Lee<sup>a,1</sup>, Changhui Lee<sup>b,1</sup>, Jaehyeok Ryu<sup>c,1</sup>, Geunpil An<sup>d</sup>, Joonsuk Bae<sup>e</sup>, Yuseul Bae<sup>f</sup>, Jeongsu Bok<sup>g</sup>, Yun Eo<sup>h</sup>, Wooseok Ham<sup>i</sup>, Yoonha Hong<sup>j</sup>, Manoj Jadhav<sup>k</sup>, Seo Yun Jang<sup>l</sup>, Jinryong Jeong<sup>l</sup>, Hyon-Suk Jo<sup>l</sup>, Sylvester Joosten<sup>l</sup>, Beomkyu Kim<sup>l</sup>, Bobae Kim<sup>l</sup>, Chong Kim<sup>l</sup>, Dongguk Kim<sup>l</sup>, Minsuk Kim<sup>l</sup>, Shin Hyung Kim<sup>l</sup>, Wonjun Ko<sup>l</sup>, Sehwook Lee<sup>l</sup>, Sanghoon Lim<sup>l</sup>, Jessica Metcalfe<sup>l</sup>, Zisis Papandreou<sup>l</sup>, Sangwoo Park<sup>l</sup>, Junseop Shin<sup>l</sup>, Hwidong Yoo<sup>l</sup>, Maria Zurek<sup>l</sup>

<sup>a</sup>Department of Physics, Sungkyunkwan University, Suwon 16419, Republic of Korea

<sup>b</sup>Department of Physics, Kyungpook National University, Daegu 41566, Republic of Korea

<sup>c</sup>Department of Physics, Pusan National University, Busan 46241, Republic of Korea

<sup>d</sup>Department of Physics, Kangwon National University, Gangneung 25457, Republic of Korea

<sup>e</sup>Department of Physics, Yonsei University, Seoul 03722, Republic of Korea

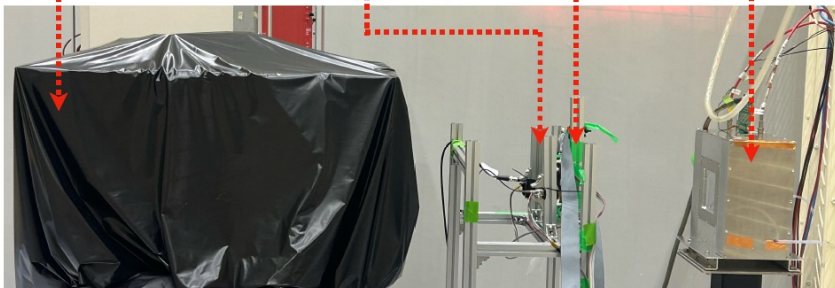
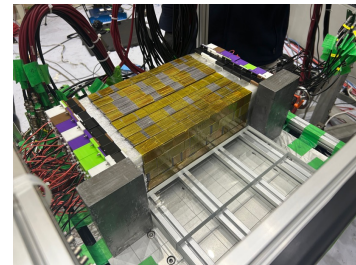
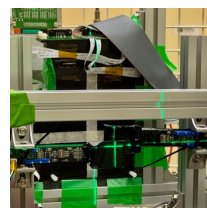
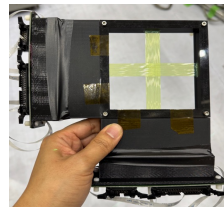
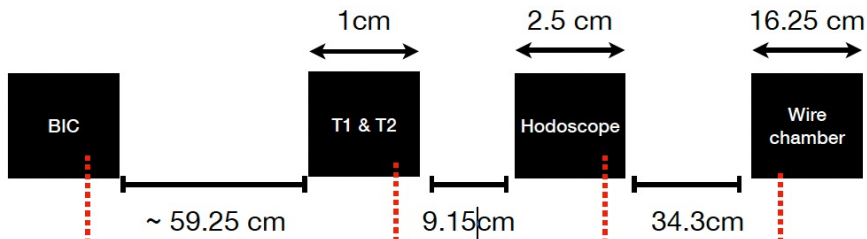
<sup>f</sup>Argonne National Laboratory, Lemont, IL 60439, U.S.A.

<sup>g</sup>Department of Physics, University of Regina, Regina, Saskatchewan, S4S 0A2, Canada



[et] 24 Apr 2026

# Second beam test at KEK PF-AR in Mar 2025

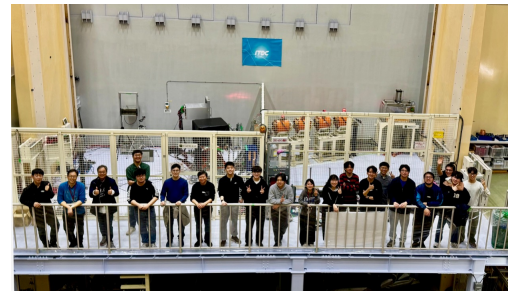
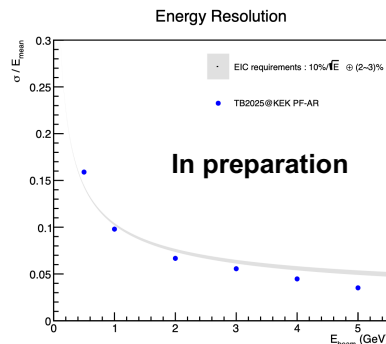
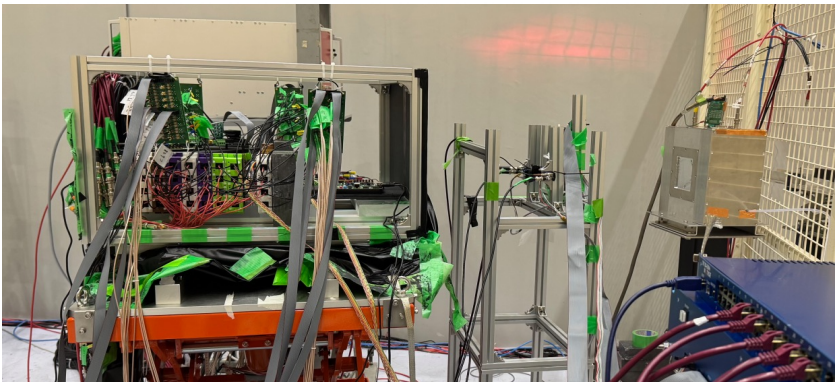


Bigger, Deeper Pb/SciFi setup: 32x12x21cm<sup>3</sup>

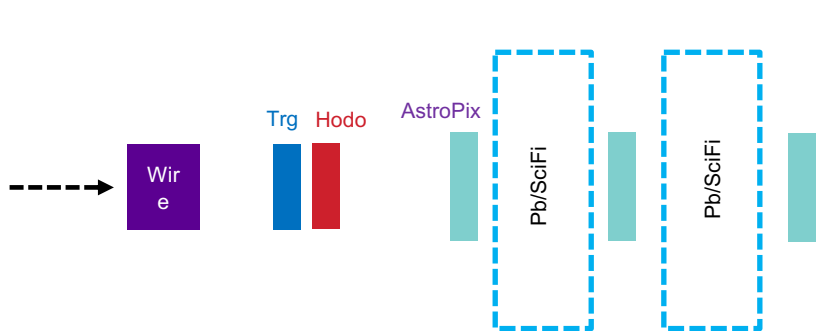
Refine calibration procedure

Smooth operation with new DAQ

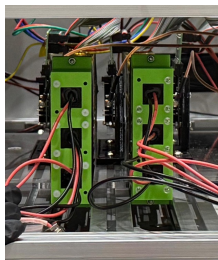
Operation of auxiliary detectors: Hodoscope, Trigger



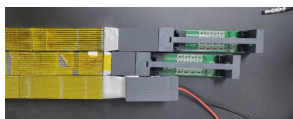
# Third beam test at KEK PF-AR in June 2025



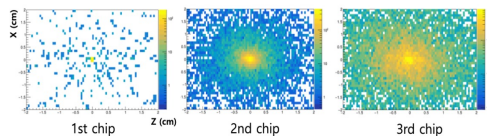
3 AstroPix Layers to be synchronized



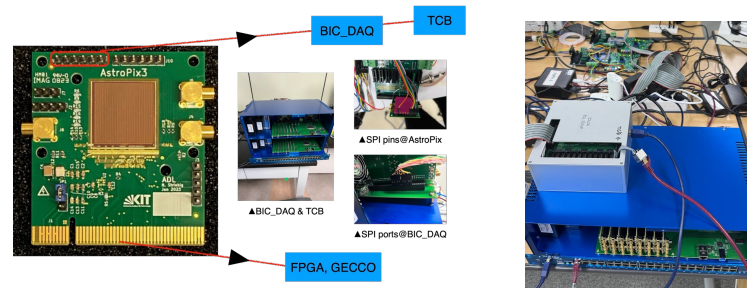
3 AstroPix Layers between Pb/SciFi



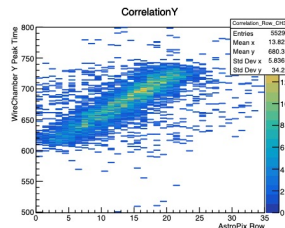
S13,S14 SiPM with LG



Hit distribution from electron beam in GEANT4 simulation



Synchronization between AstroPix and other detectors in the DAQ system produced in Korea  
Preparation of the shower profile study.  
SiPM and Light Guide test



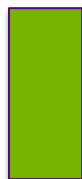
correlation of beam position at AstroPix and drift time at Wire Chamber



# Recent beam test at CERN PS T10 in July 2025

Beam ( $e, \pi, \mu$ )  
Maximum  $10^6$ /spill

Cherenkov



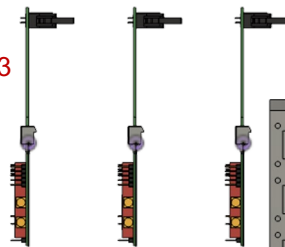
Trigger  
(15x15mm<sup>2</sup>)



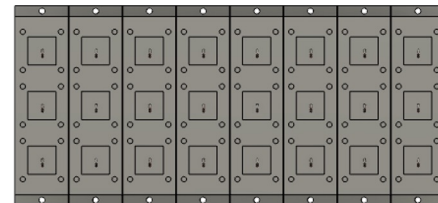
Hodoscope



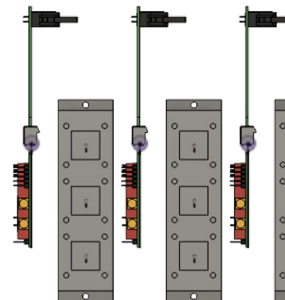
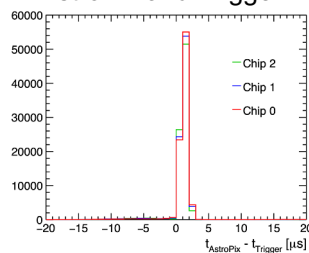
AstroPix v3



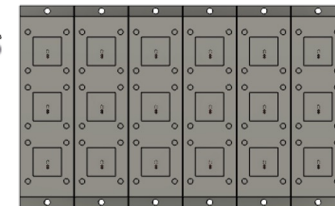
(A) AstroPix layers



Time between  
AstroPix and Trigger



(B) AstroPix between Pb/SciFi

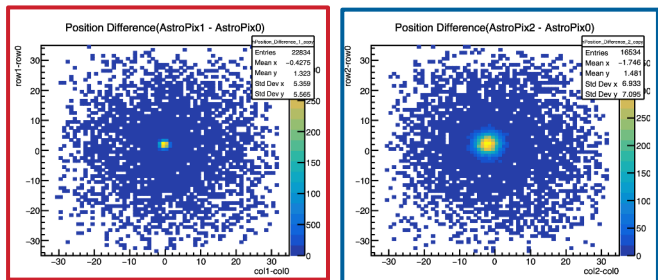


Goal: Proof-of-principle imaging of electrons and pions using a synchronized setup of AstroPix in the beam environment, using synchronized data taking between AstroPix and Pb/SciFi



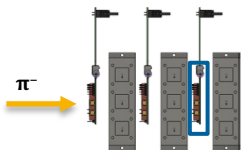
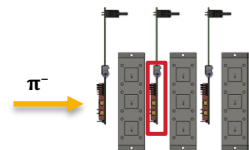
# Different shower profiles of electrons and pions

## Pions

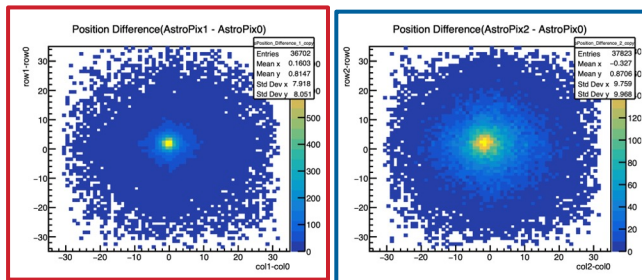


Position difference between 1st-2nd layer

Position difference between 1st-3rd layer

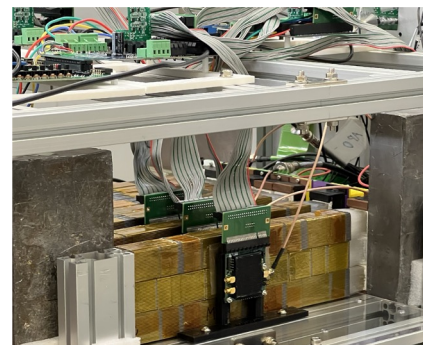
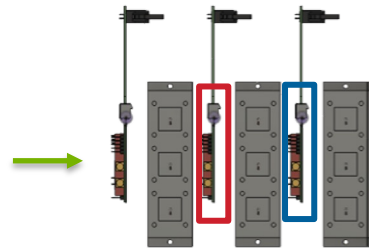
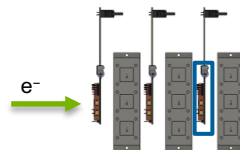
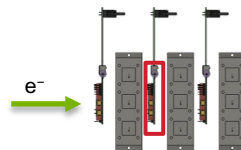


## Electrons



Position difference between 1st-2nd layer

Position difference between 1st-3rd layer



Proof-of-principle imaging of electrons and pions was performed using a synchronized AstroPix setup in the beam environment.

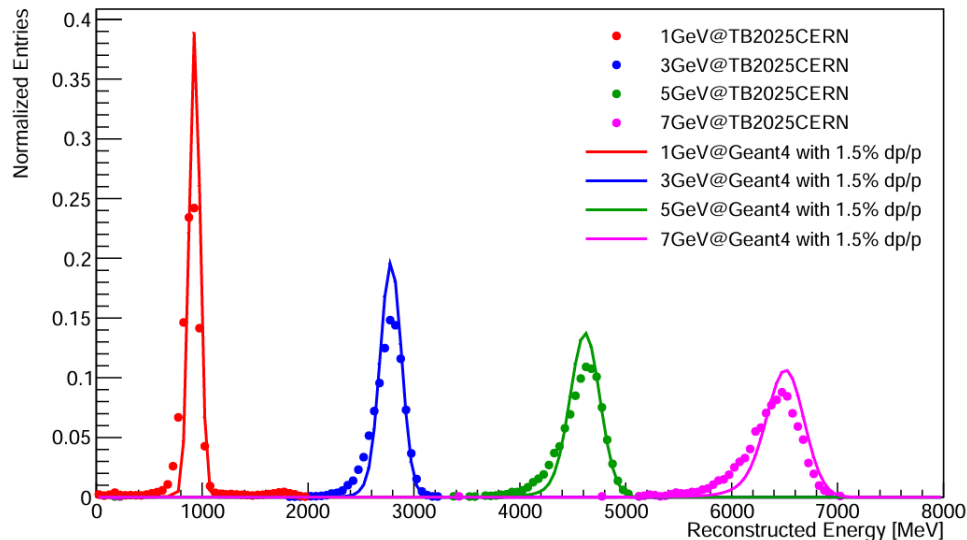
Test-Beam Performance of the AstroPix Silicon Sensor for Imaging Calorimetry

Yoonha Hong<sup>c1</sup>, Jeongsu Bok<sup>c1</sup>, Geunpil An<sup>d</sup>, Joonsuk Bae<sup>e</sup>, Yunsuek Bae<sup>f</sup>, Regina Caputo<sup>g</sup>, Yun Eo<sup>h</sup>, Woosock Ham<sup>a</sup>, Woohyeon Heo<sup>b</sup>, Manoj Jadhav<sup>g</sup>, Seo Yun Jang<sup>g</sup>, Jinryong Jeong<sup>g</sup>, Hyon-Suk Jo<sup>b</sup>, Sylvester Joosten<sup>b</sup>, Beomkyu Kim<sup>a</sup>, Bobae Kim<sup>b</sup>, Chong Kim<sup>a</sup>, Dongguk Kim<sup>a</sup>, Minsuk Kim<sup>d</sup>, Shin Hyung Kim<sup>b</sup>, Woojong Kim<sup>a</sup>, Wonjun Ko<sup>h</sup>, Changhui Lee<sup>b</sup>, Hyungjun Lee<sup>a</sup>, Jason Sang Hun Lee<sup>g</sup>, Jongwon Lee<sup>g</sup>, Kyeongpil Lee<sup>g</sup>, Schwook Lee<sup>b</sup>, Sangwoo Park<sup>a</sup>, Jaehyeok Ryu<sup>h</sup>, Bogyeon Seo<sup>h</sup>, Jessica Metcalfe<sup>b</sup>, Minsub Shim<sup>h</sup>, Junseop Shin<sup>h</sup>, Hwidong Yoo<sup>g</sup>, Maria Zurek<sup>b</sup>, Sanghoon Lim<sup>g</sup>

arXiv:2605.07681

<sup>a</sup>Department of Physics, Sungkyunkwan University, Suwon 16419, Republic of Korea  
<sup>b</sup>Department of Physics, Kyungpook National University, Daegu 41566, Republic of Korea  
<sup>c</sup>Department of Physics, Pusan National University, Busan 46241, Republic of Korea  
<sup>d</sup>Department of Physics, Kangwon National University, Gangneung 25457, Republic of Korea  
<sup>e</sup>Department of Physics, Yonsei University, Seoul 03722, Republic of Korea  
<sup>f</sup>Department of Physics, Chung-Ang University, Seoul 06974, Republic of Korea  
<sup>g</sup>Department of Physics, University of Seoul, Seoul 02504, Republic of Korea  
<sup>h</sup>Argonne National Laboratory, Lemont, IL 60439, U.S.A.  
<sup>i</sup>NASA Goddard Space Flight Center, Greenbelt, MD 20771, U.S.A.

# Electron energy measurement



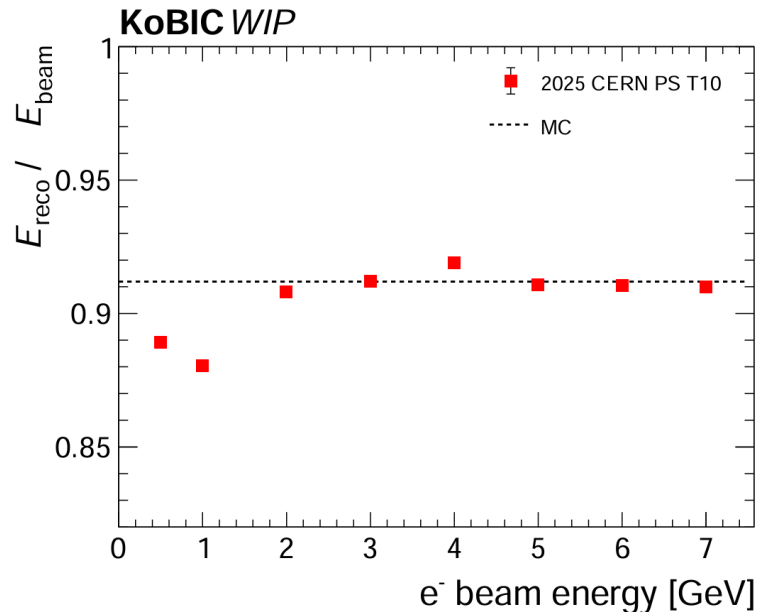
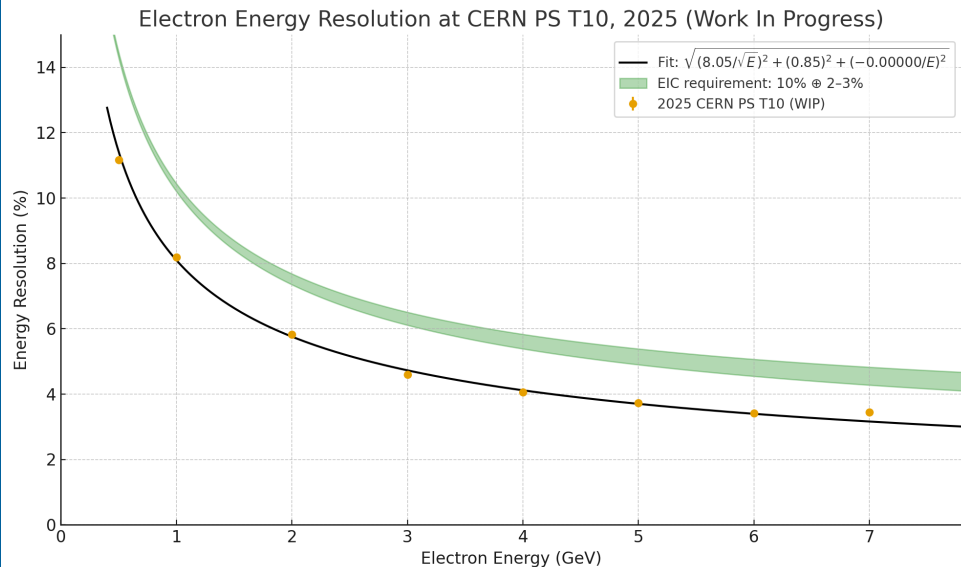
Applying beam momentum spread  
(dp/p) to simulation

$$\sigma_{\text{total}} = \sqrt{\sigma_{\text{original}}^2 + \sigma_{\text{spread}}^2}$$

\*1.5% dp/p : Expected momentum spread for the 3 mm collimator setup

- According to the GEANT4 simulation, energy resolution is affected by the beam momentum spread.
- At CERN PS T10, the beam momentum spread is 0.6~15%, depending on the collimator setup. (arXiv:2507.02567)
- With an additional momentum spread of 1.5%, the resolution in the simulation is shown with the data.

# Energy Resolution & Linearity



**Resolution :** (Stochastic) 8.05 %, (Constant) 0.85 %, (Noise) 0.00 %

$$\frac{\sigma_E}{E} = \frac{a}{\sqrt{E}} \oplus b \oplus c/E$$

- The energy resolution has been progressively improved, achieving the best performance in this beam test.

# EIC Science Program in First Years of Operations

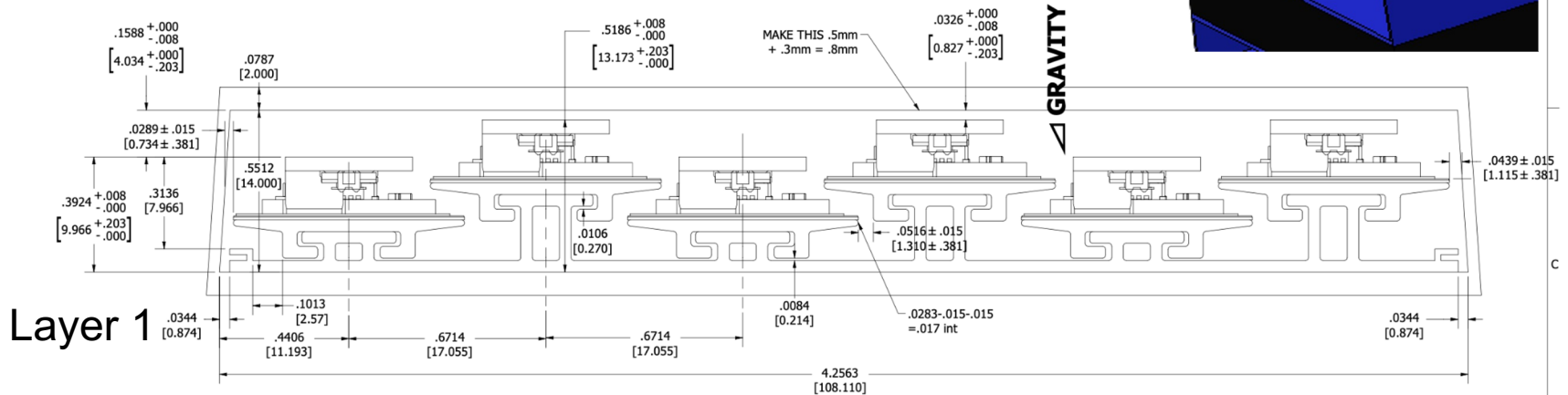
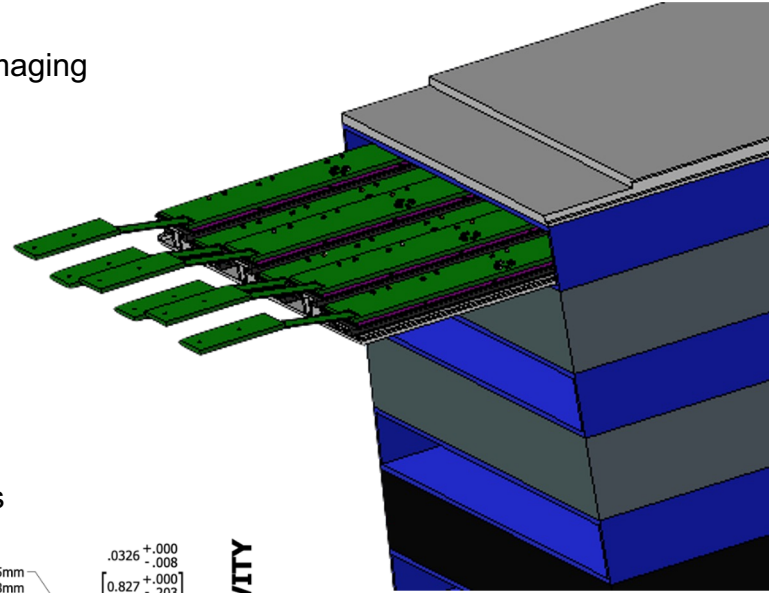
Year - 1	Year - 2	Year - 3	Year - 4	Year - 5
Start with Phase 1 EIC	Phase 1 EIC + long. electron polarization	Phase 1 EIC + long. electron polarization + proton polarization	Phase 1 EIC + long. electron polarization + proton polarization + spin rotators	Phase 1 EIC + long. electron polarization + proton polarization + operation of hadron spin rotators
<b>New Capability:</b> Commission electron polarization in parallel	<b>New Capability:</b> Commission proton polarization in parallel	<b>New Capability:</b> Commission running with hadron spin rotators	<b>New Capability:</b> Commission hadron accelerator to operate with not centered orbits	
<b>Run:</b> 9 GeV electrons on 115 GeV/u heavy ion beams (Ag, Nb, Cu, Ru)	<b>Run:</b> 9 GeV long. polarized electrons on 130 GeV/u Deuterium	<b>Run:</b> 9 GeV long. polarized electrons on 130 GeV transverse polarized protons	<b>Run:</b> 9 GeV long. polarized electrons on 110 GeV Au	<b>Run:</b> 9 GeV long. polarized electrons on 110 GeV Au
<b>Physics:</b> Dense Gluon Systems QCD in Nuclei	<b>Physics:</b> QCD in Nuclei Flavor structure of nucleons and nuclei	<b>Physics:</b> Flavor structure of nucleons and nuclei	<b>Physics:</b> Dense Gluon Systems QCD in Nuclei	<b>Physics:</b> Dense Gluon Systems QCD in Nuclei
	<b>Run:</b> Last weeks: 9 GeV long. polarized electrons and 130 GeV polarized protons	<b>Run:</b> Last weeks switch to longitudinal proton polarization	<b>Run:</b> 9 GeV long. polarized electrons on 275 GeV transverse and longitudinal polarized protons	<b>Run:</b> 9 GeV long. polarized electrons on 166 GeV transverse and longitudinal polarized He-3
	<b>Physics:</b> Flavor structure of nucleons and nuclei	<b>Physics:</b> Spin	<b>Physics:</b> Spin	<b>Physics:</b> Spin

**Backup**



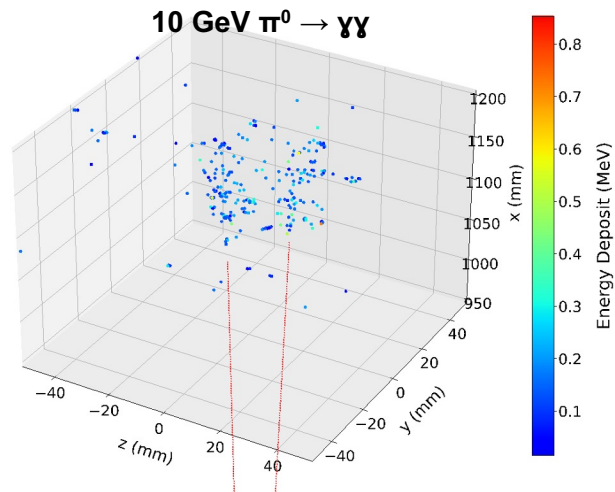
# BIC Imaging Layers

- **A Module** is an electrically testable elementary unit for imaging layers
  - **9 AstroPix HV-CMOS chips** - 2 cm × 2 cm
  - Base Plate (Aluminum) slides on Stave rail
  - Rigid(-flex) PCB readout
- 12 modules form a **Stave** and readout at the end of the Sector using End of Tray Card (FPGA) (~217cm)
- **Tray** consists of (6, 7, 7, 8) Staves (56 Staves/Sector)
- There are a total of 48 sectors with a length of ~435 cm
- The BIC is made up of 31104 Modules using ~280k chips

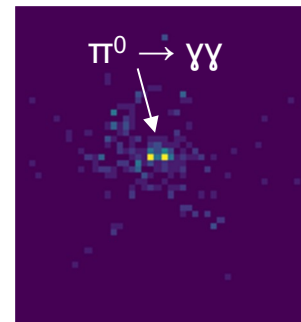
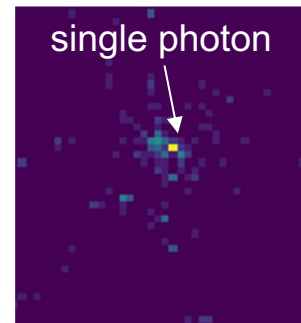
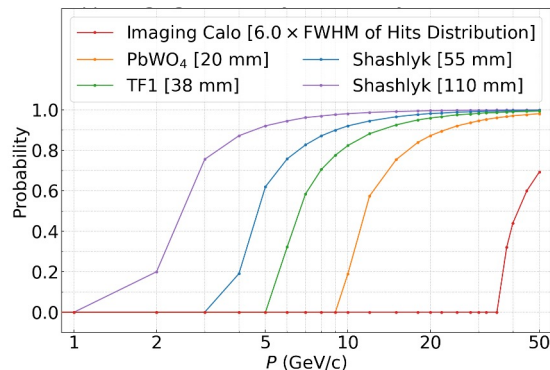


# Particle Identification

## Photons and Neutral Pions



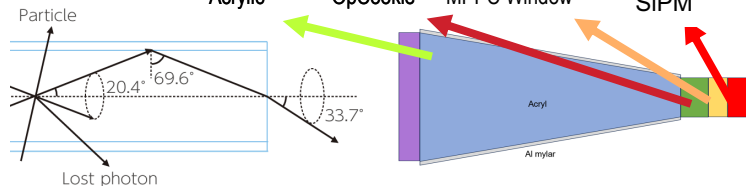
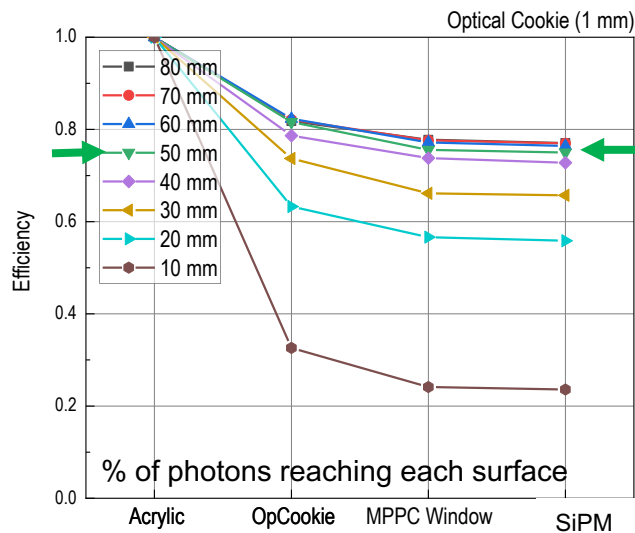
Merging probability  $\gamma/\pi^0$  (upper limit)



- Precise position resolution and shower imaging allow for excellent separation of  $\gamma/\pi^0$  based on the 3D shower profile
- Upper limit anticipated from the **AstroPix position resolution and shower profile: well above 10 GeV**
- First insights from a simple **neural network approach** (~82% pion rejection at 90% efficiency above 10 GeV with the current status of model training - lower limit)

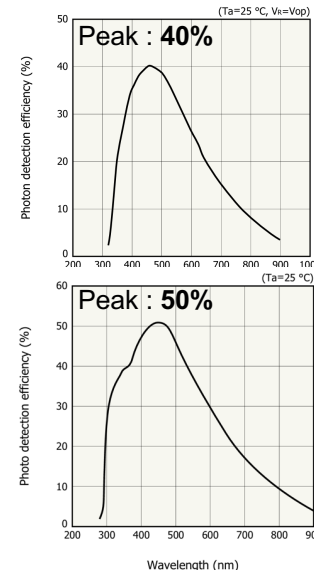
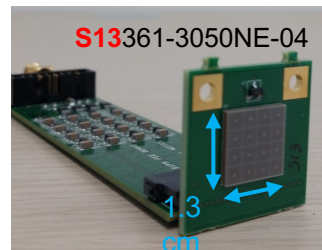
# Development of the readout system

Transmission efficiency with LG length



Optical simulation with random photon generation ( $\phi < 20.4^\circ$ )  
**Efficiency drops rapidly with lightguide length under 5 cm.**

Silicon Photomultiplier



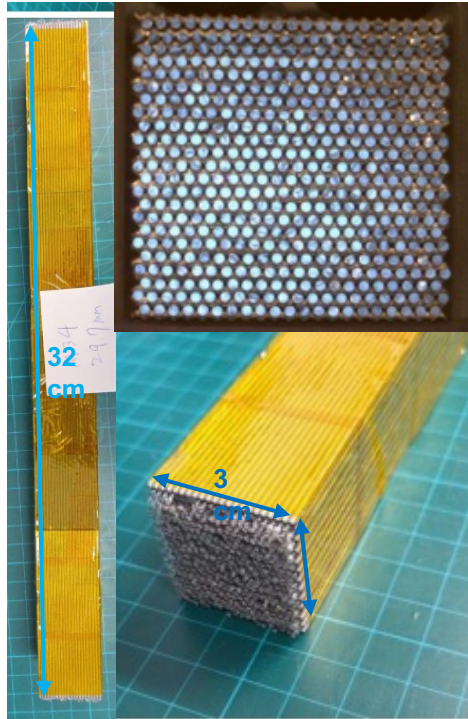
**SiPM (S13 & S14, Hamamatsu)**

**Photon detection efficiency (S14 > S13)**

**Dark count rate (S14 > S13)**

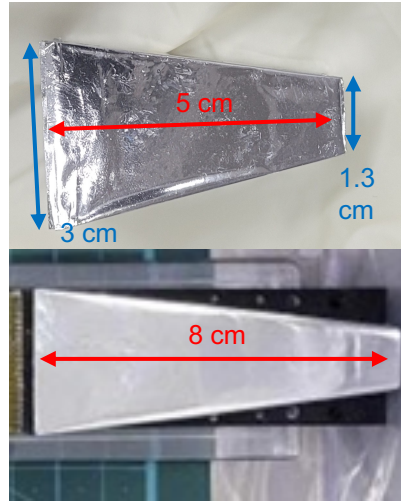
**Neutron radiation hardness (S13 > S14)**

# Development of the readout system



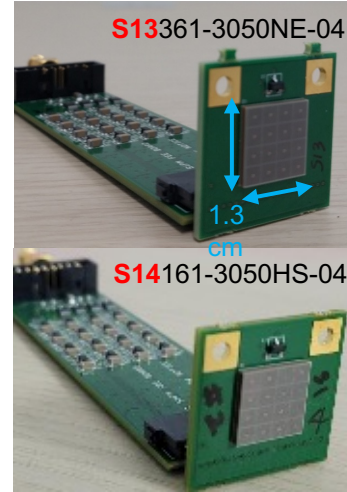
## 1. Calorimeter

Pb plates & Scintillating fibers



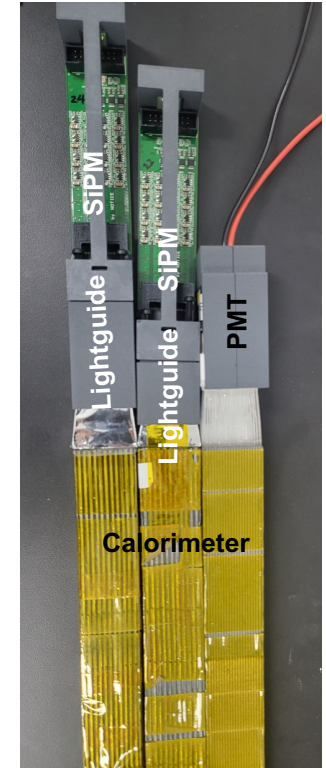
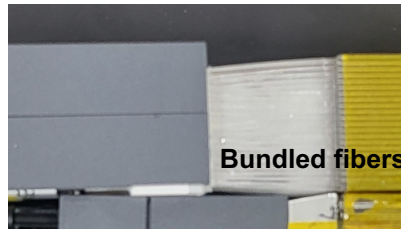
## 2. Lightguide

Square, acrylic lightguides wrapped with Al Mylar (5, 8 cm)  
Bundled fiber for PMT

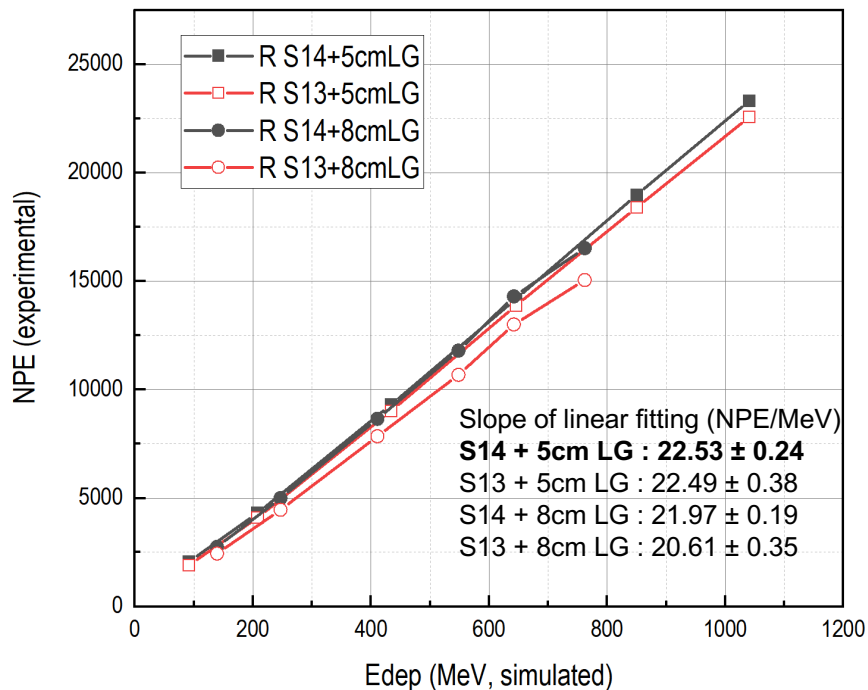


## 3. Photon sensor (Hamamatsu)

SiPM (S13 & S14)  
PMT

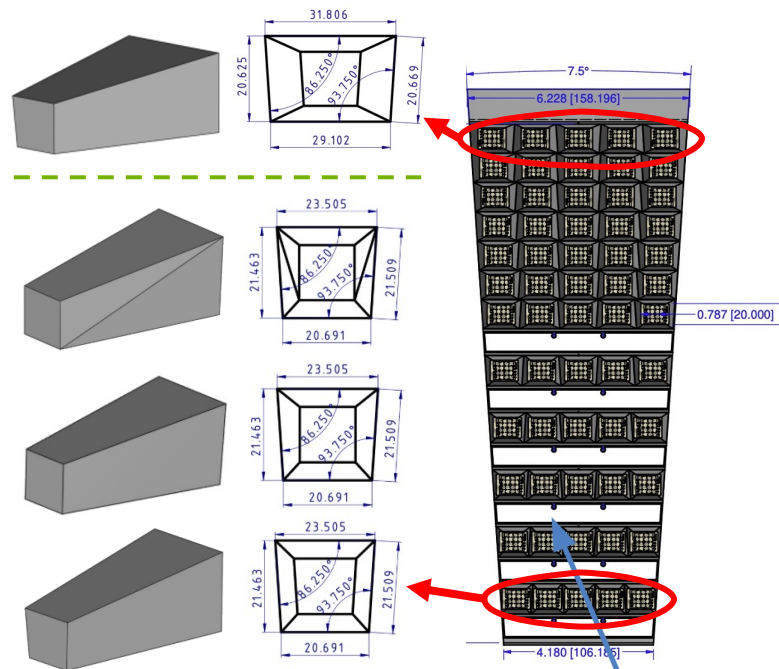


# Development of the readout system



Beam test results

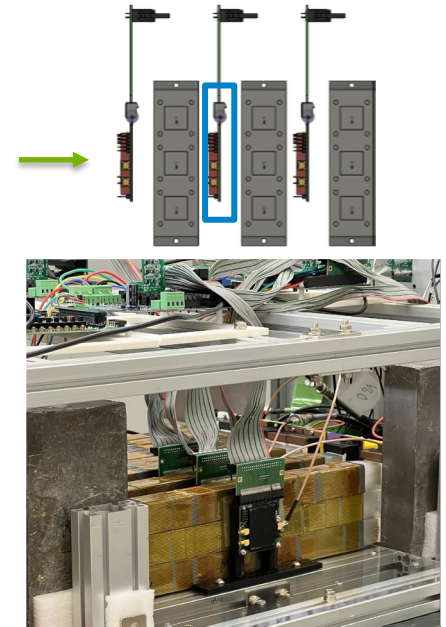
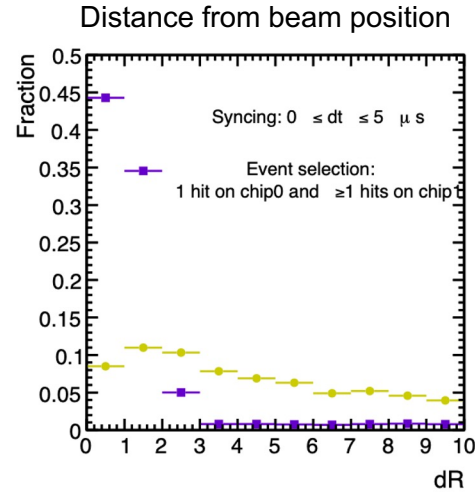
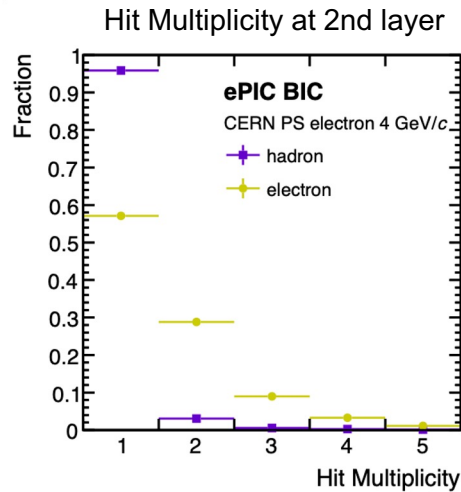
1. Linearity of SiPM readout
2. Small difference in light yield for SiPM readout (< 10%)



Future plan

Lightguides matching the dimensions for actual deployment will be tested

# Different shower profiles of electrons and pions



- The hit multiplicity and distance from the beam position at the second AstroPix v3 chip differ for electrons and pions.
- The electron/pion separation power in the beam test condition will be enhanced by AstroPix information on shower shape.

# Pb/SciFi Prototype Module Production

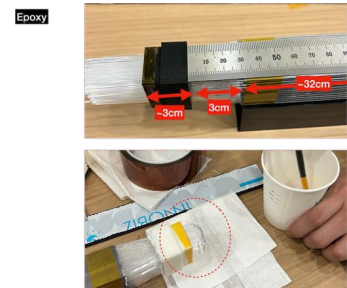
1) Pb plate preparation



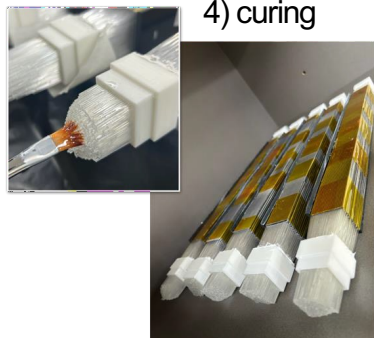
2) Stacking with fiber



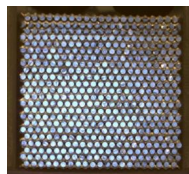
3) Cutting fiber



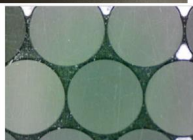
4) curing



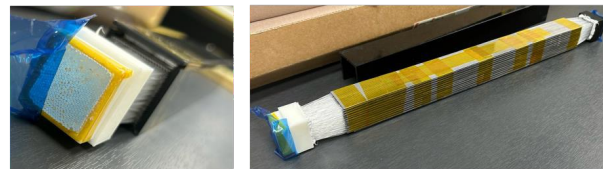
Produced 33 of 32x3x3 cm<sup>3</sup> unit modules for beam test



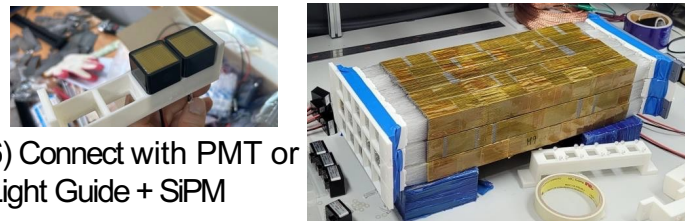
5) polishing



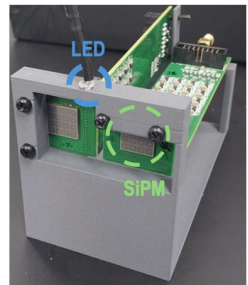
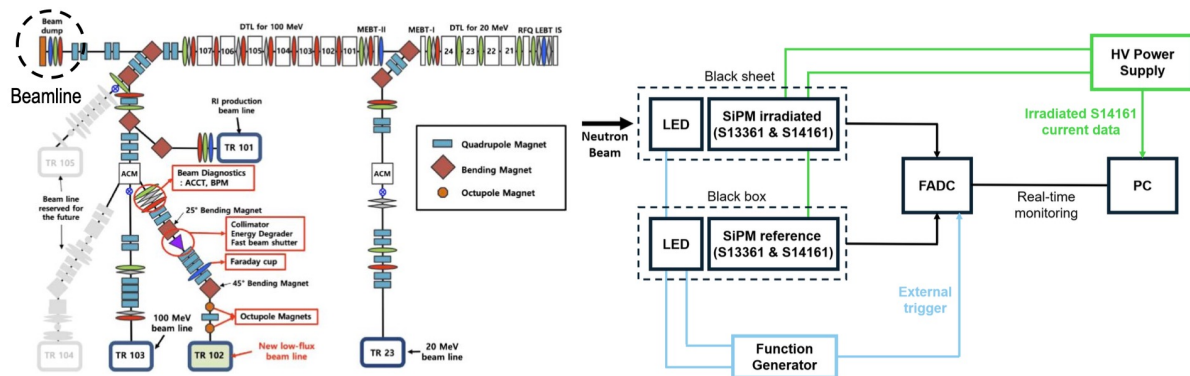
원장본



6) Connect with PMT or Light Guide + SiPM



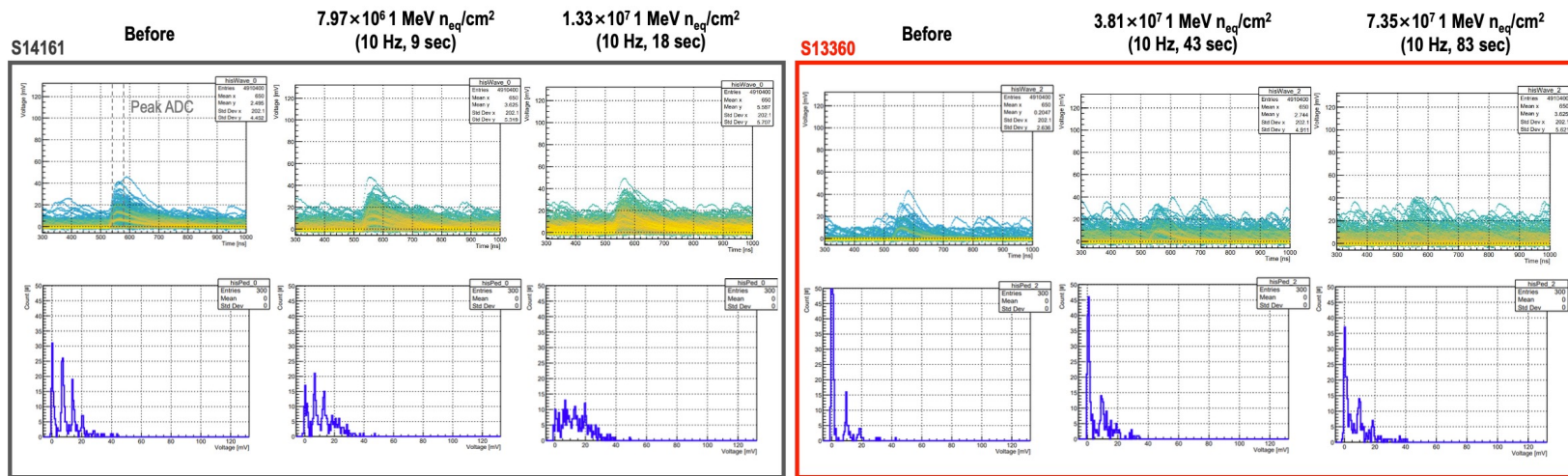
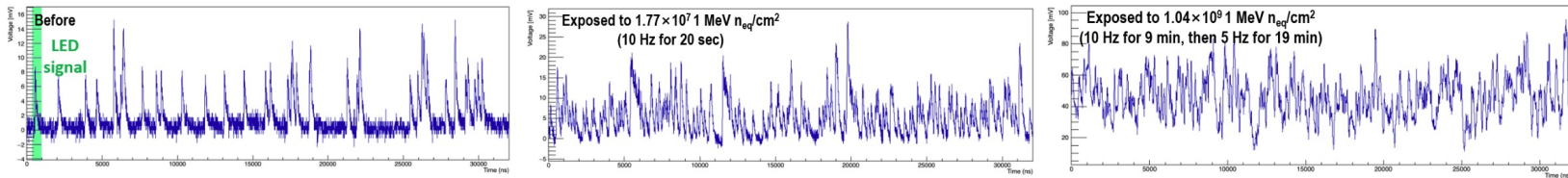
# SiPM radiation hardness test



Date	Beam setting	Fluence [1MeV $n_{eq}/cm^2$ ]	LED Hz	Remarks
Jun 26	10 Hz × 5 min 10 Hz × 4 min 5 Hz × 19 min	$1.04 \times 10^9$	100 Hz	Comparison of S13 and S14 S14: higher photon detection efficiency S13: lower dark count rate
Sep 5	1 Hz × 60 min vs 5 Hz × 12 min	$6.20 \times 10^8$ (each)	100 Hz vs 500 Hz	Two pairs (S14 & S13) Same fluence, test for rate dependence

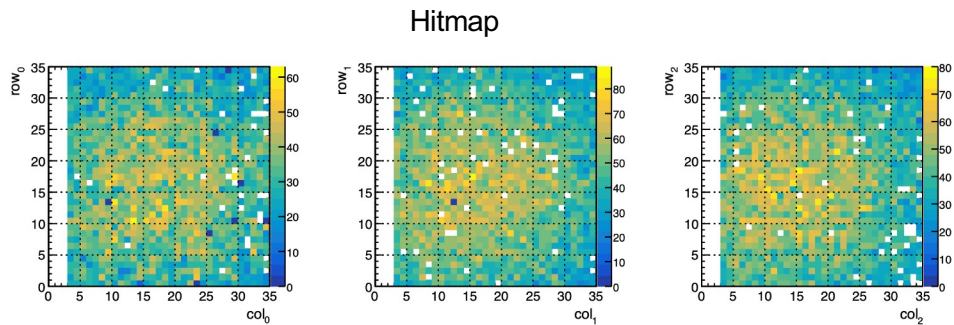
- Real-time SiPM irradiation monitoring: Two SiPMs (Hamamatsu S13361 and S14161) were irradiated with neutron beams at KOMAC, and monitoring of single-photoelectron response (LED window), dark current rate, and baseline shifts.
- Performance change evaluation: Pre- and post-irradiation IV measurements were performed to quantify leakage-current increases and breakdown-voltage shifts for both devices.

# SiPM radiation hardness test

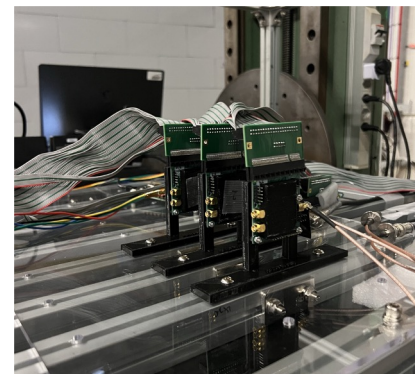
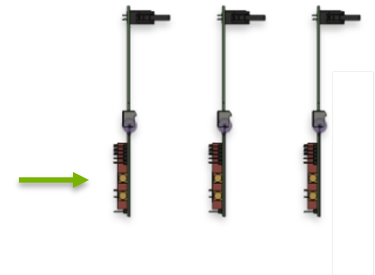
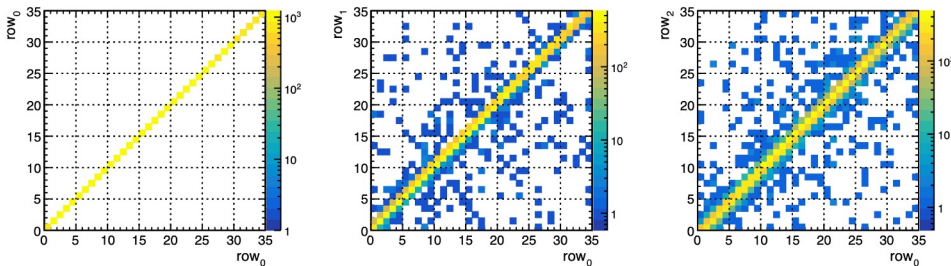


S14161 shows a lower ADC sigma as a noise indicator, while S13361 exhibits lower leakage current and better single-photoelectron peak separation during irradiation.

# Synchronization of three layers of AstroPix v3

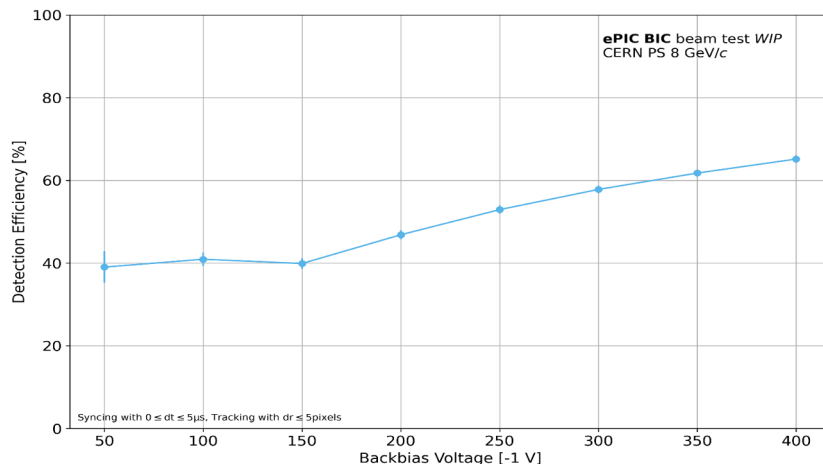


Correlation to the upstream chip

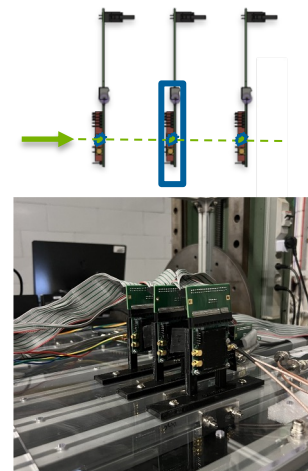


- Three AstroPix v3 layers are tested using the electron/pion beam at CERN PS.
- Correlations to the upstream show good alignment and data synchronization between AstroPix.
- The integrated system works well at a 2.7 kHz trigger rate on the beam. (BIC rate < 1 kHz/chip)

# AstroPix v3 chip tracking efficiency (Shallow depletion depth)



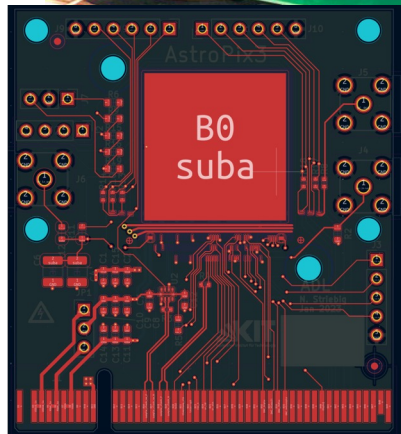
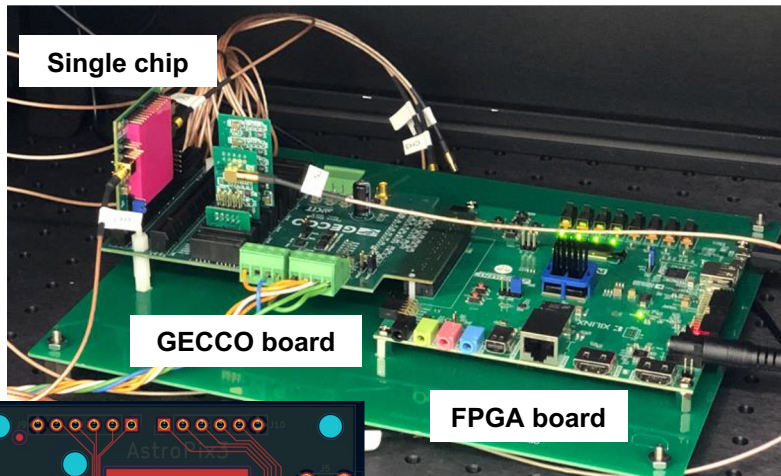
Detection efficiency in percentage as a function of bias voltage



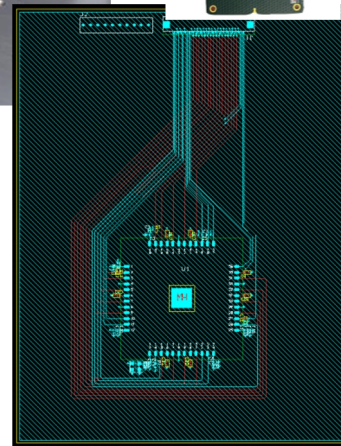
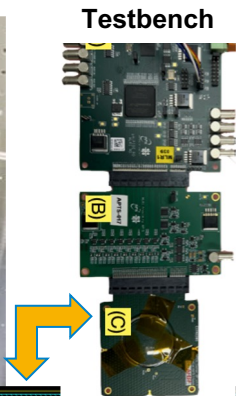
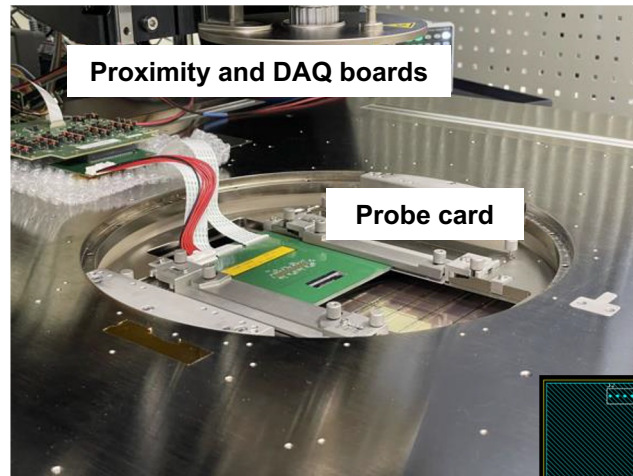
- Proof-of-principle for a method to measure tracking efficiency of AstroPix v3 under beam conditions.
- It is calculated as the probability of having a hit on the position estimated from hits on the first & third chips.
- The shallow depletion of v3 was known and is addressed in v5 with increased depletion.

# AstroPix test system

## AstroPix v3 lab measurement setup



## ALICE ITS3 ER1 wafer probing system



- Initial version of probe card for AstroPix v3: a simple version for the chip carrier board.
- Uses already validated GECCO and FPGA board setup, which will be connected with flexible cables
- Recently, the probe card design for v3 has been done by a local manufacturer in Korea

# AstroPix v5 Specifications

## Pixel Matrix:

- 36 cols x 34 rows
- 32 Columns with Standard NMOS Comparator
- 2 Columns with dynamic Feedback
- 1 Column with NMOS Comparator and Resistor Load
- 1 Column with NMOS Comparator and PMOS Load

500u Pixel Pitch and 300u Pixel Size  
3 Tunebits per Pixel

Pixel Dynamic Range 20 keV - 700 keV

Noise Floor 5 keV (2% @ 662 keV)

Bias Voltage up to 400-500 V to maximize depletion

Fully NMOS Comparator

In Pixel amplifier with Dynamic Feedback option for improved Dynamic Range (2 columns)

## Power Consumption:

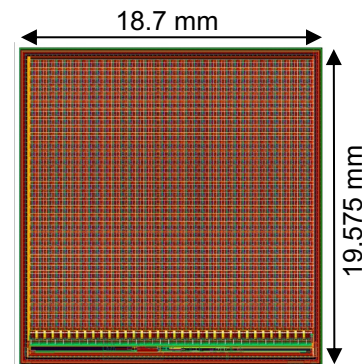
Pixel 4.6  $\mu$ W

Pixel matrix 5.3 mW

Digital 2.2 mW

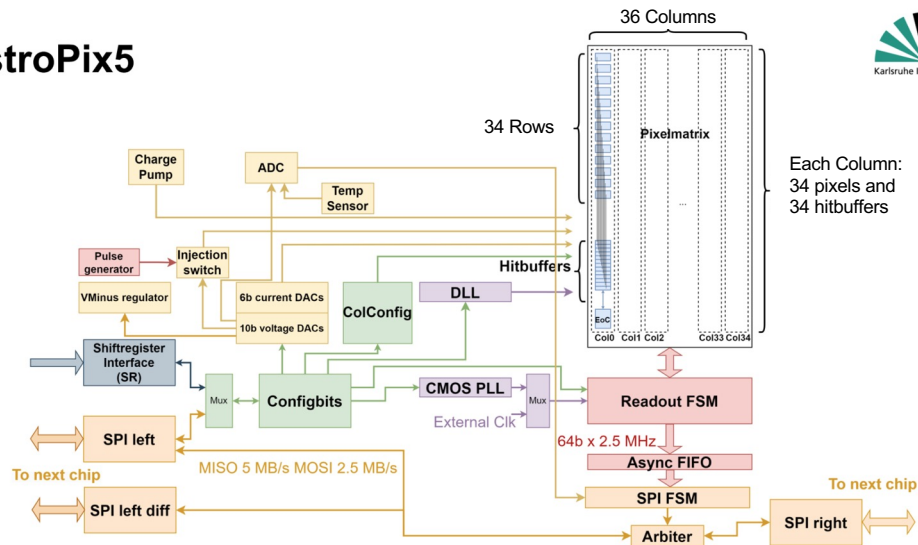
– 700  $\mu$ W DigitalTop

**Total:**  $\sim$ 2 mW/cm<sup>2</sup>



18.7 mm  $\times$  19.575 mm chip

## AstroPix5



# Relevant AstroPix v6 Specifications

## Pixel Matrix:

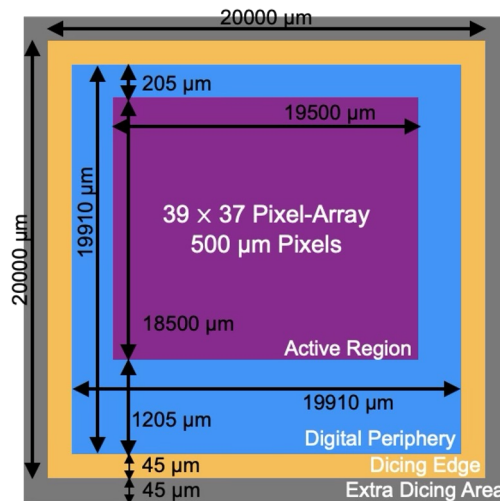
- 39 cols x 37 rows
- 35 Columns with Standard NMOS Comparator
- 2 Columns with dynamic Feedback
- 1 Column with NMOS Comparator and Resistor Load
- 1 Column with NMOS Comparator and PMOS Load
- 500u Pixel Pitch and 300u Pixel Size
- 3 Tunebits per Pixel
- Pixel Dynamic Range 20 keV - 700 keV
- Noise Floor 5 keV (2% @ 662 keV)
- Bias Voltage up to 400-500 V to maximize depletion
- Fully NMOS Comparator
- In Pixel amplifier with Dynamic Feedback option for improved Dynamic Range (decision from v5 validation)

## Power Consumption:

- Pixel 4.6 uW
- Pixel matrix 5.3 mW
- Digital 2.2 mW
- 700 uW DigitalTop

**Total:** ~2 mW/cm<sup>2</sup> for 2x2 cm<sup>2</sup> chip

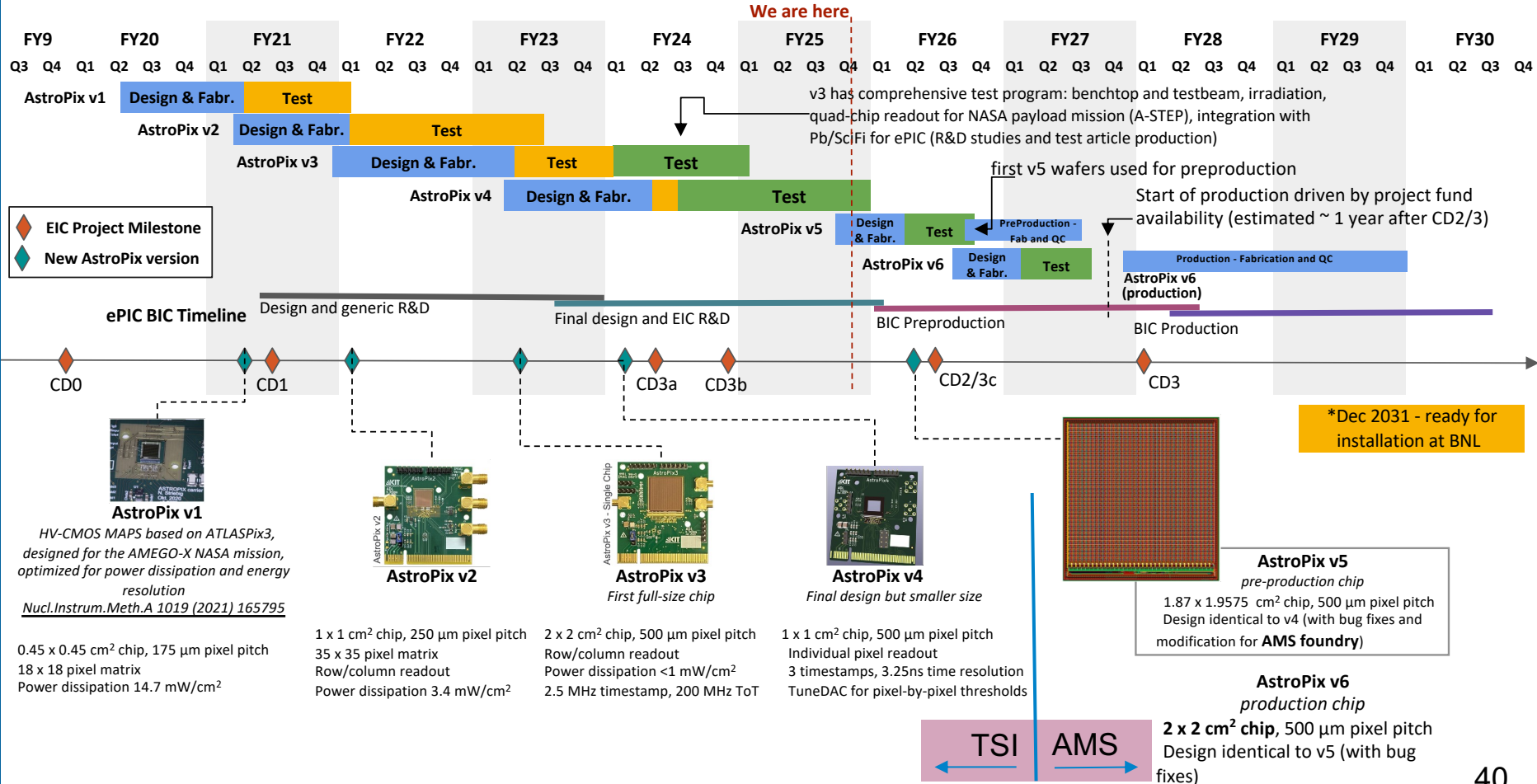
AstroPix v6 will be scaled up to **20 mm × 20 mm** chip size



AstroPix v6 do not expect major design changes. Scaled up chip size to required 2 cm × 2 cm dimension with bug fixes and selection of dynamics range feedback and NMOS comparator.

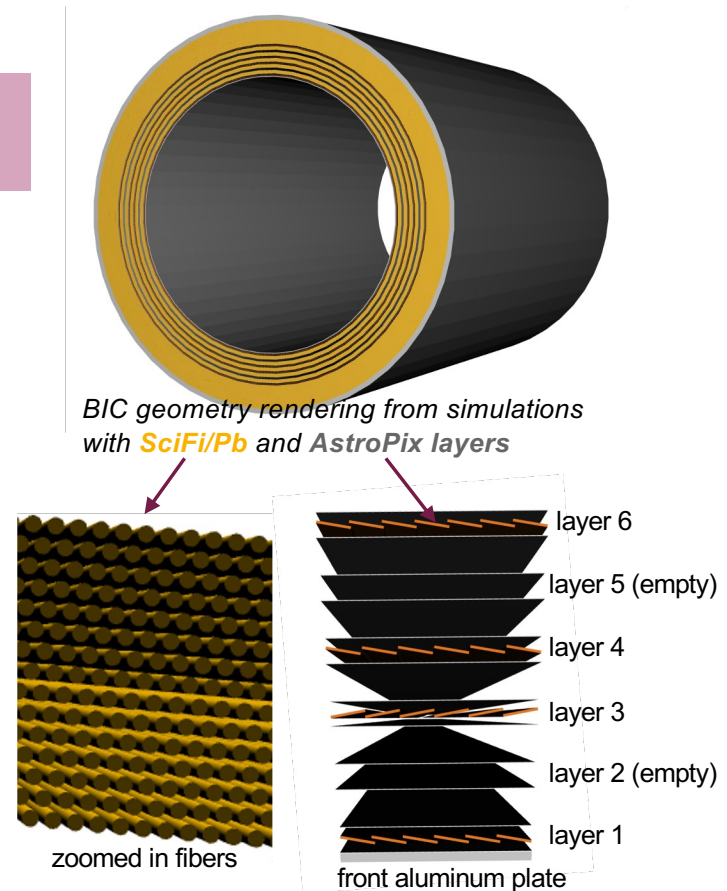
# AstroPix Timeline

◆ Not shown: Early CD4 (Oct 2032)  
◆ Critical Design Review (CDR)  
◆ Program Implementation Review (PIR)  
◆ CD4 (Oct 2034)



# Simulations

- **Simulation** results based on realistic detailed simulations benchmarked in beam and bench environment.
- **Official ePIC geometry implementation** with realistic detailed BIC geometry implemented
  - Full detector with all subsystems, the majority of services
  - Cladded SciFi placed in Pb and glue mixture
  - AstroPix layers: “turbofanned” staves with chips; expected dead areas included
- **Realistic digitization and reconstruction** with effective model for the optical output including:
  - **light attenuation and phe statistics** (benchmarked in GlueX and adjusted for fiber differences)
  - **light guide efficiencies** from optical simulations
  - thresholds based on **SiPM HGCROC-template noise**
- EM and Had responses in **simulations benchmarked with the beam test data up to 10 GeV**

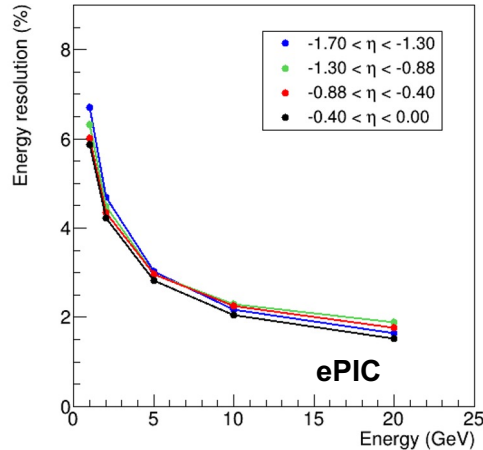


# Energy Response

## Photons

- Energy resolution meets requirement  $< 10\%/\sqrt{(E)} + (2-3)\%$
- Benchmarked in beam tests.

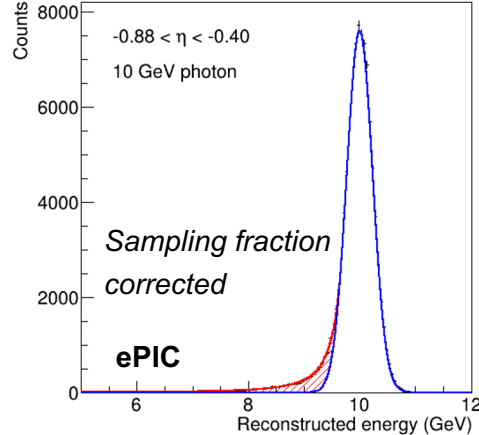
### Energy Resolution



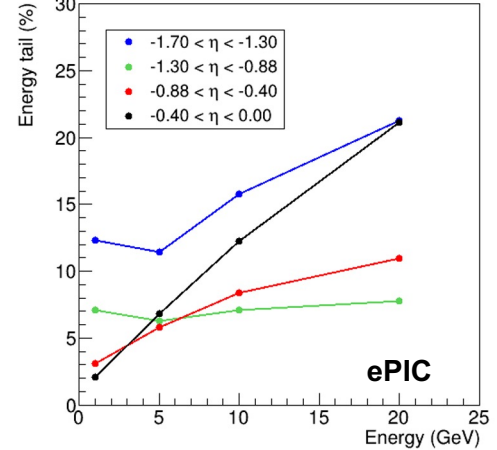
### Fit parameters

$\eta$	$a/\sqrt{(E)}$ [%]	$b$ [%]
(-1.7, -1.3)	6.60(0.03)	0.66(0.04)
(-1.3, -0.88)	6.11(0.01)	1.24(0.01)
(-0.88, -0.4)	5.91(0.02)	1.24(0.02)
(-0.4, 0)	5.85(0.01)	0.88(0.02)

### Energy Deposit Shape



### Energy Response Tail

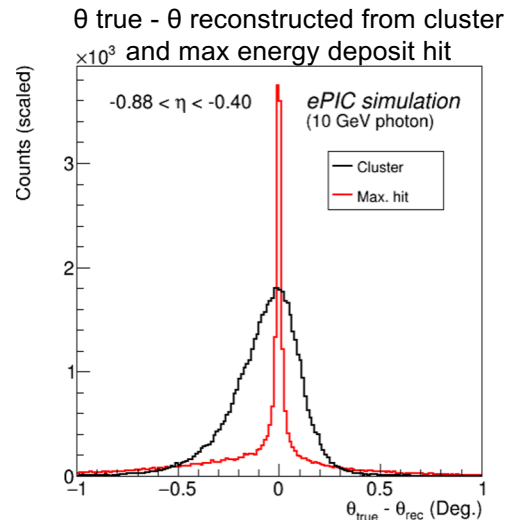
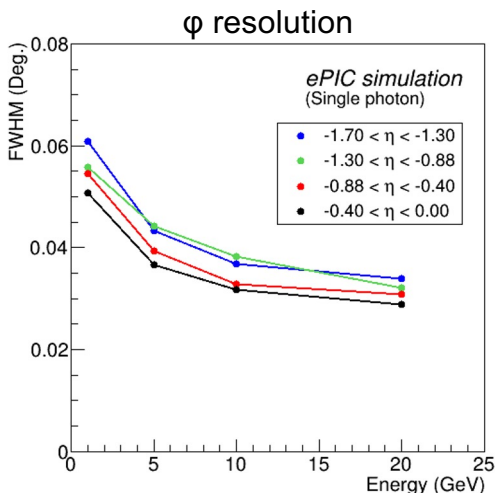
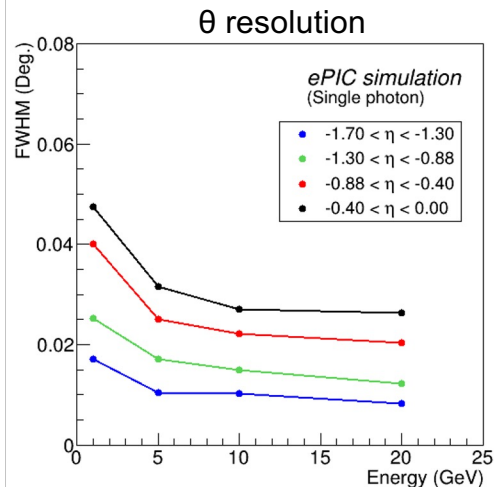


- Based on the realistic simulation of the Pb/ScFi part of the calorimeter with attenuation and photoelectron statistics
- Resolution extracted from a Crystal Ball fit  $\sigma$ 
  - Energy tail evaluated as a difference between the Crystal Ball and the extrapolated Gaussian fit

# Position Response with AstroPix

## Photons

- Angular resolution well below 0.1 deg (on the level of single pixel resolution) well below any tower-like calorimetry
- Position measurements benchmarked in beam tests
- **Methods:** max-deposit hit at the shower-start layer compared to  $\theta$ ,  $\phi$  from 3D topological clustering.  
**Result:** hit-level method gives an order of magnitude better angular resolution.
- Small dependence seen with changing  $\eta$ . The  $\phi$  resolution is worse than  $\theta$  because the shower particles are smeared by the magnetic field.



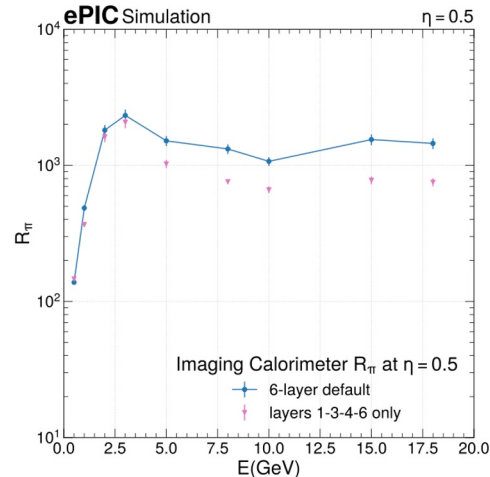
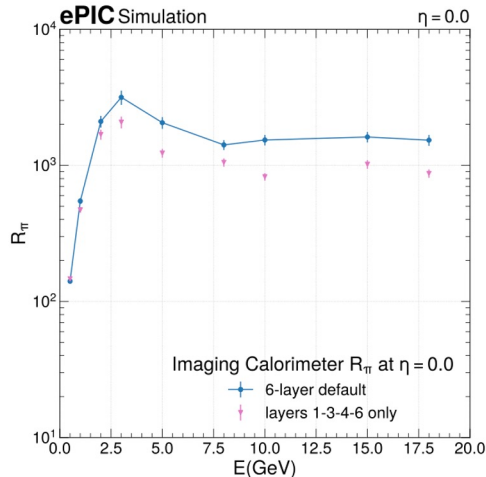
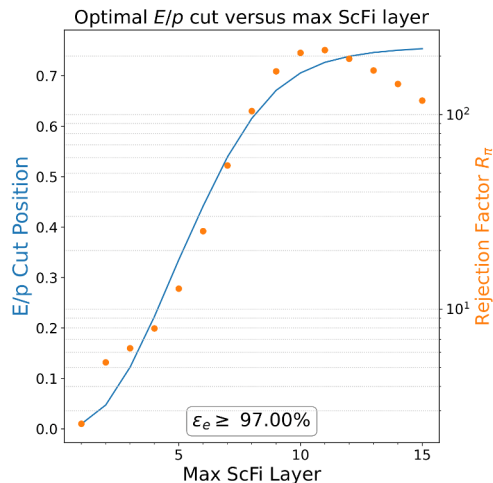
# Particle Identification

## Electrons and Pions

- High  $e/\pi$  separation down to 1 GeV/c delivers required  $\sim 90\%$  electron purity with combined ECAL+DIRC. See *backup for more details*. (P-DET-ECAL-BAR.26.10.05)

**Method: Optimized cut on  $E/p$**  from different depth of Pb/ScFi layers at very high electron efficiency + **Convolutional neural network** utilizing energy and spatial information for shower

Example for 2 GeV  $e/\pi$



Scan through different number of SciFi/Pb layers to find the optimal maximal depth.  
**Peak ~ layer 10-11**

**Pion rejection:** 4-layer achieves  $>10^3$  at low-mid energies where it is most needed; 6-layer achieves  $>10^3$  across most energies and is a good upgrade option for high-luminosity, higher-energy runs.

# Sectors and Mechanics

Design constraints defined and mostly frozen; geometries and envelopes documented; BIC and Project engineers aligned.

**Design challenges:** self-supporting sector with ePIC integration; FEA in progress using delamination data at CF-Pb and Pb/SciFi interfaces.

## Sector test articles:

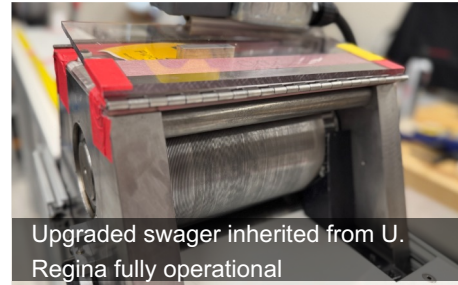
- 0.5 m GlueX-style matrix reproduction completed, critical equipment commissioned. SciFi QC procedure established.
- Short bulk Pb/SciFi with GlueX fibers and CF-frame integration ongoing
- Long 3 m article next; full-scale 4.35 m sector for preproduction targeted for 2026

**AstroPix tracker mechanics:** staves integrated into trays → modules slide onto trays → trays slide into sector drawer/slot.

- First extruded-aluminum mechanical article built and tested for tolerances and stability



*First tracker test article - module base with stave like railing*



Mechanical risks are under active test with a defined path to full-scale sector in 2026. First test articles of tracker support and SciFi matrix tested.

# Wafers and Modules

## Chip testing scope

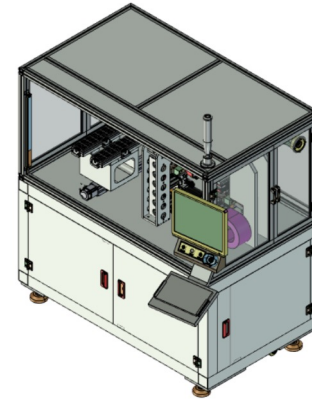
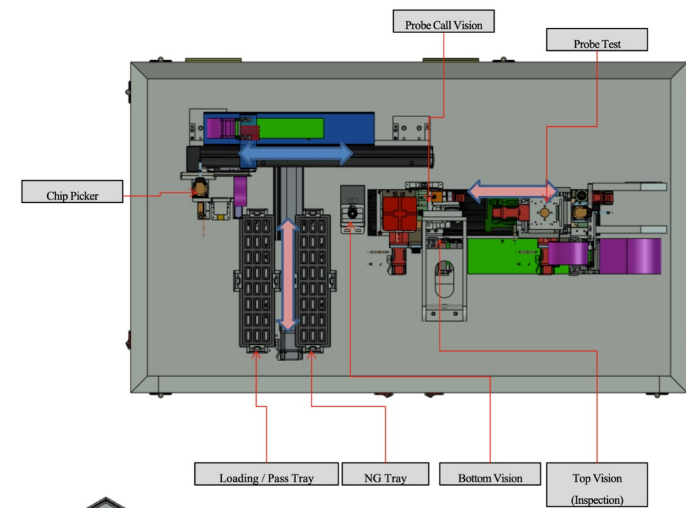
- **Scale:** 31,104 modules, 279,936 AstroPix chips.
- **Need:** automated chip-level testing before module assembly with established QC procedures.  
**QC coverage:** sensor bias with HV, analog and digital checks, injection scans for relative energy calibration.

## v3 Pilot (this fall → 2025 Q4)

- **Probe card:** simple v3 card for the carrier board, built with a Korean manufacturer; uses validated GECCO + FPGA setup; leverages ALICE experience
- **Throughput target:** validate the QC flow on ~80 v3 chips this fall
- **Automation:** single-chip v3 test machine from C-ON Tech due 2025 Q4 to run initial validation, refine the test flow, document the full chain, and deliver qualified v3 chips for prototype modules

## Preproduction (v5/v6, FY26-FY27)

- **Design update:** v5 ready in FY26 for preproduction
- **Process readiness:** finalize procedures and acceptance criteria



*Single-chip v3 test machine is under production by C-ON tech in Korea*

We are moving from a validated v3 pilot to a preproduction workflow for v5/v6, establishing throughput, traceability, and acceptance criteria needed to de-risk delivery of 280k+ chips for 31k modules.

# Wafers and Modules

**Scope:** single-flavor module for more than 30,000 units

**AstroLinx (PCB):** optimized for module performance, meets required operation specs; first design completed and reviewed. Expected this fall; another iteration anticipated.

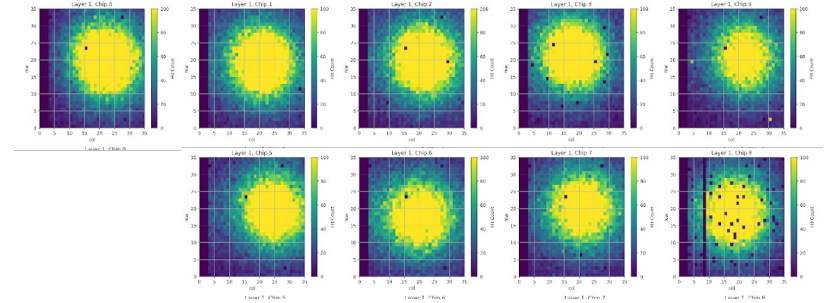
**9-chip PCB test article:** In hand to mock up AstroLinx, verify power-distribution stability and daisy-chain readout. Successfully tested on a bench with sources.

**Module builds:** plan to assemble about six prototypes this fall using an extruded aluminum base, AstroPix v3, and AstroLinx; test PCB designed to validate electrical performance; initial base-plate prototype produced to define assembly steps.

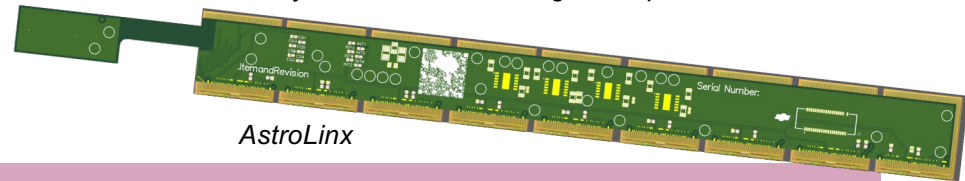
**Open design items:** base-plate locking mechanism, adhesive selection, flex-PCB connector finalization



*AstroPix v3 9-chip board*



*Daisy chain readout through 9 chips with source scan*



*AstroLinx*

AstroLinx has been reviewed and the 9-chip PCB validated with sources; we are proceeding to module test articles to lock performance and assembly while closing the few remaining design choices for preproduction.

# ESB (End of Sector Box) and DAQ

**Function:** on-detector power distribution and data collection.  
Encloses cooling and provides mechanical protection.

**Interfaces with ePIC:** integration procedure nearly final; service requirements provided (electrical cabling, cooling, DAQ).

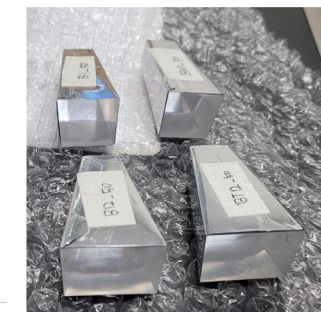
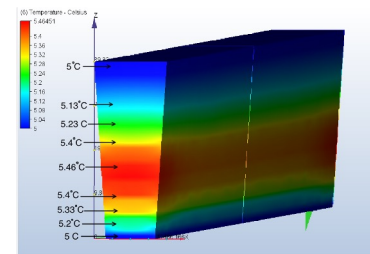
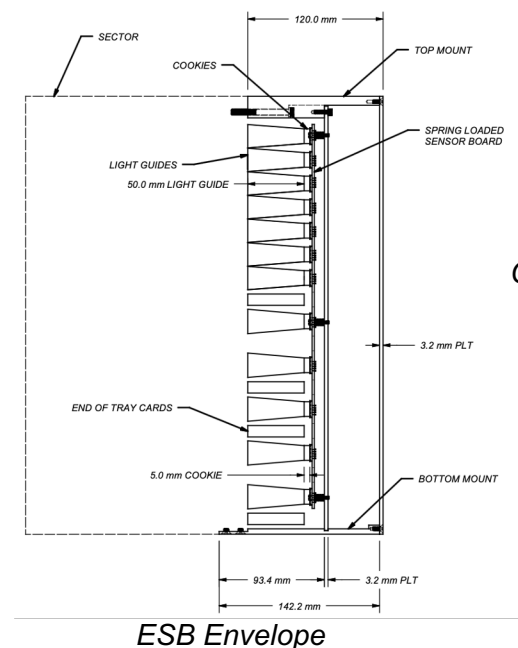
**Thermal/mechanical validation: Ongoing:** development and construction of PCB (thermal) test article to validate CFD calculations.

**SiPM readout path:** CALOROC → common RDO → FELIX.

- Lightguides design complete, geometry fixed.
- CALOROC design under submission; evolution of HGCROC with CMS and ALICE heritage.
- SiPMs passed FDR at CD-3a/3b.

**AstroPix readout path:** End-of-Tray Card (ETC) → FELIX.

- ETC to be designed by NASA GSFC; start pending contract processing; leverages existing AstroPix FPGA board designs.



ESB envelope and interfaces nearly finalized; CALOROC at submission; ETC awaits contract start, with existing designs reducing risk. SiPM design completed. Lightguides and cookies nearly finalized.

# System Testing

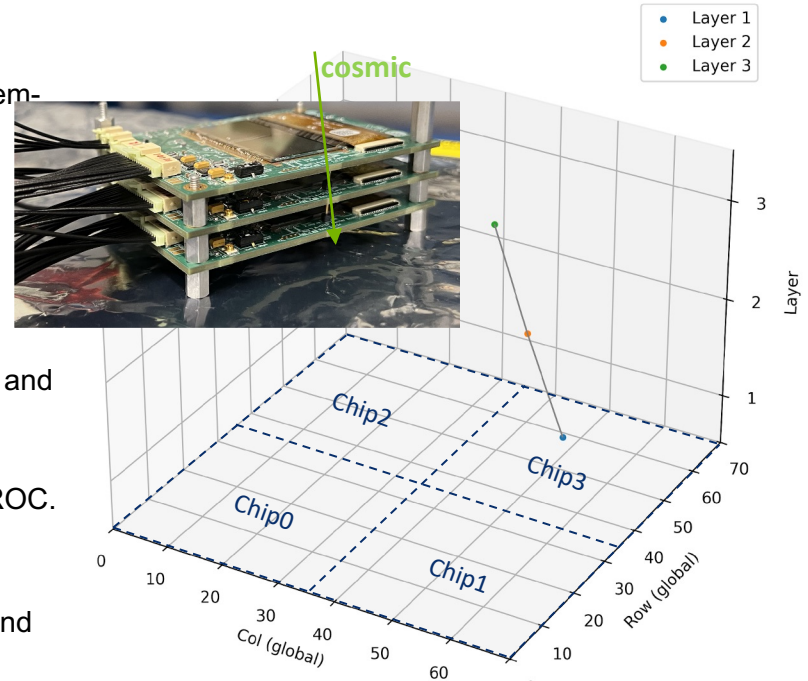
**Function:** validate system-level performance and design, establish system-level QC/QA procedures.

**AstroPix imaging layers:** multichip and multilayer test articles to demonstrate daisy-chaining, synchronization, and energy, position, and timing response; bench and beam tests completed.

**Pb/SciFi calorimetry:** Bulk ( $\sim 15.5 X_0$ ) and SFILs ( $\sim 1.5 X_0$ ) tested with  $e$  and  $\pi$  beams, validating energy response and  $e/\pi$  separation in the Pb/SciFi section; simulations benchmarked.

- **ANL-built SFILs:** first units show good MIP response with HGCROC.

**Large-scale integration (PED, PREP):** readout and large-test-article integration underway to validate system-level performance, calibration, and QC/QA at the AstroPix Tray and Sector levels; outputs will also inform component-level QC procedures.



*Example cosmic event display from 3 layers of daisy-chained AstroPix v3 quad chips*

The testing program de-risks the design and defines system QC/QA for production; current validations confirm calorimeter energy performance and rate capability, and validate the simulation against data.

# Production Plan

\*iCRADA still to be signed

vendor  
site

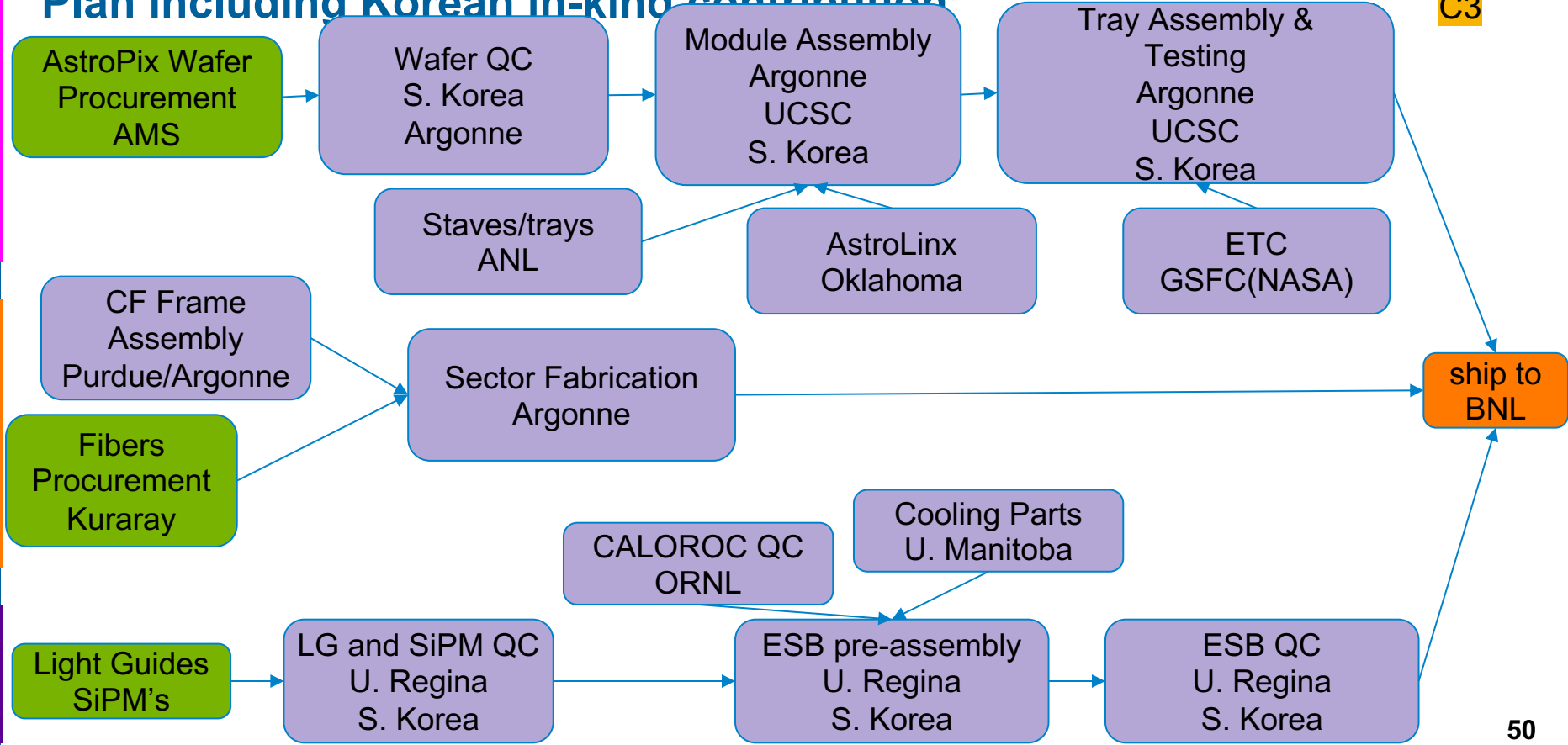
## Plan including Korean in-kind contribution\*

C3

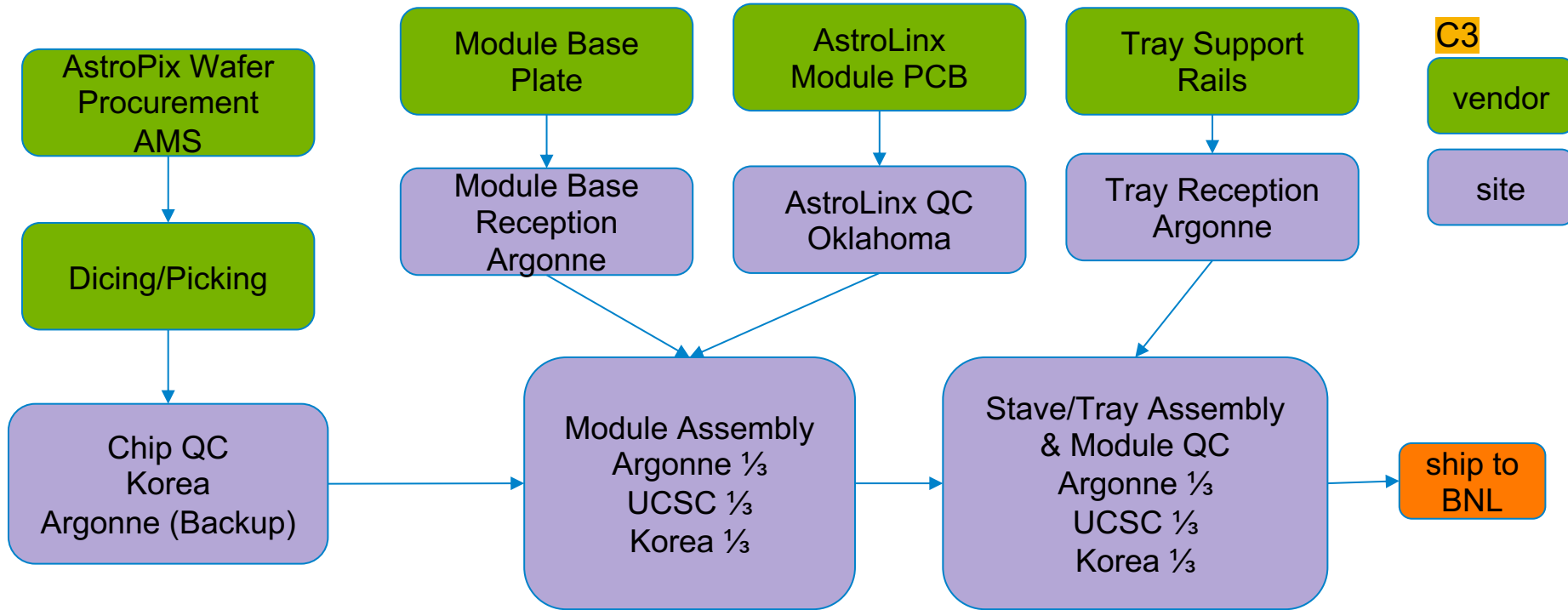
Tracker

Sector

ESB



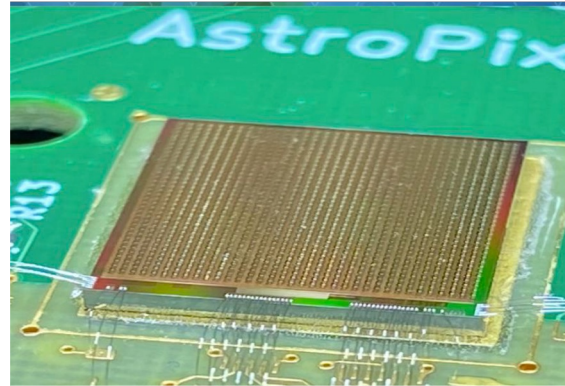
# Workflow: Tracker Layers



- Wafer procurement with AMS
- One flavor for each part
- Thorough QC at bare chip level to ensure good chips/sensors
- Keep all steps as simple and streamlined as possible for industrial style manufacturing

# Wafer QC Probing

1. Purchase wafers
2. Dicing
  - a. picked by vendor, comes in a waffle pack
3. Automated wafer (or chip) QC probing
  - a. Visual inspection
  - b. Full chip functionality
  - c. Sensor IV
  - d. Pixel-by-pixel performance
4. Ship to module assembly sites

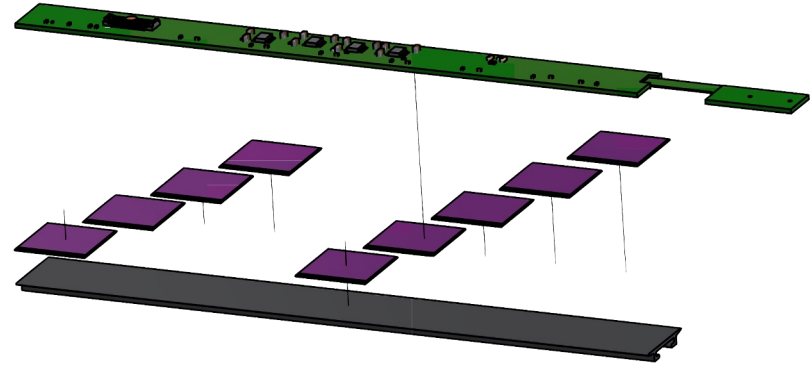


→ Expect to have a high yield after chip testing until glued on a module

Wafer QC	# parts	# per batch	# batches	hours/batch	minutes/piece	# people	Total Person hours	Inactive hours/batch	Batches/week/production line	# of production lines	Total production weeks	Total production years (w/ 85% annual efficiency)
chip probing	357,319	60	5,956	5	5	1	596	4.9	15	6	66	1.5

# Module Assembly

1. Stage chips and module base plates onto the tooling
2. Glue chips to the base plate, cure
3. Stage the module and AstroLinx onto the tooling
4. Glue AstroLinx to the module, cure
5. Wire bond
6. Pot wire bonds (under consideration)



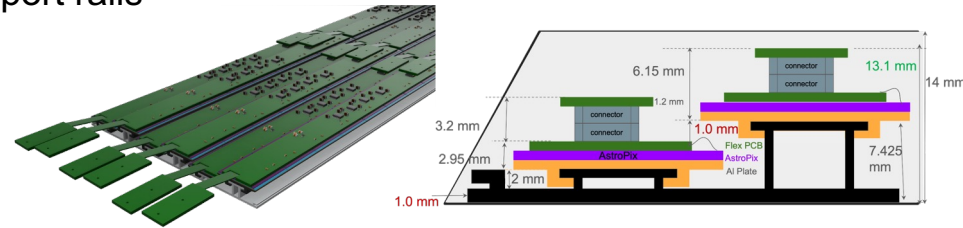
Module Assembly	# parts	# per batch	# batches	hours/batch	minutes/piece	# people	Total Person hours	Inactive hours/batch	Batches/week/production line	# of production lines	Total production weeks	Total production years (w/ 85% annual efficiency)
chip reception	303721	100	3037	0.17	0.1	1	506	0	240	3	4	0.10
module gluing	33747	6	5625	5	-	1	2812	4.5	5	18	63	1.4
gluing AstroLinx	33747	6	5625	5	-	1	2812	4.5	5	18	63	1.4
wire bonding	33747	1	33747	0.25	15	1	8437	0	160	3	70	1.60
potting	33747	6	5625	0.17	1.7	1	937	8	240	3	8	0.18

# Stave/Tray Assembly

1. Slide modules onto *the bottom row* of the tray support rails
2. plug connectors, quick test, rework
3. Slide modules onto *the top row* of the tray support rails
4. plug connectors, quick test, rework
5. Full electrical QC w/ end-of-tray card
6. Pack and ship to BNL

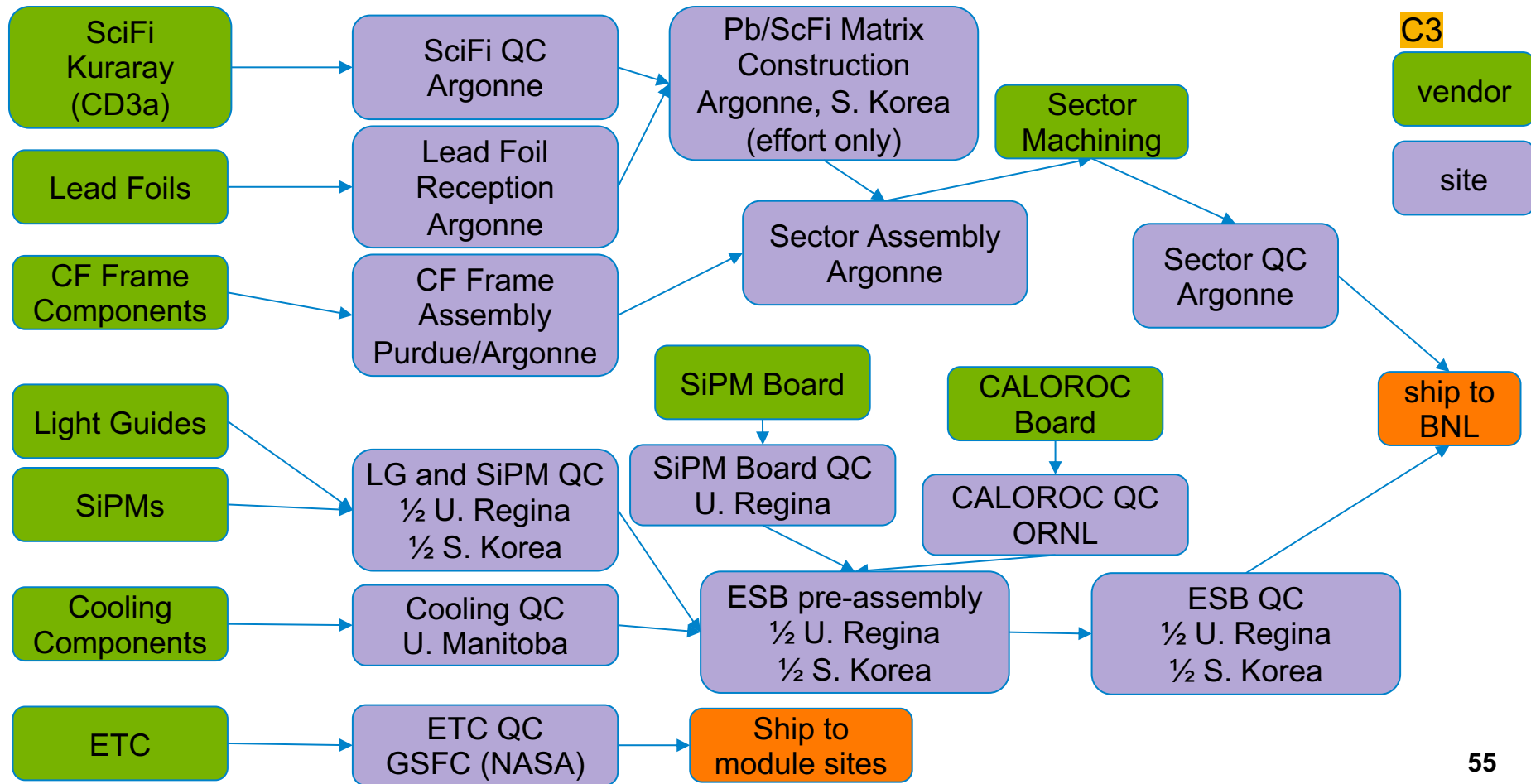
## Reworkable design:

- connectors rated for mating cycles
- no gluing

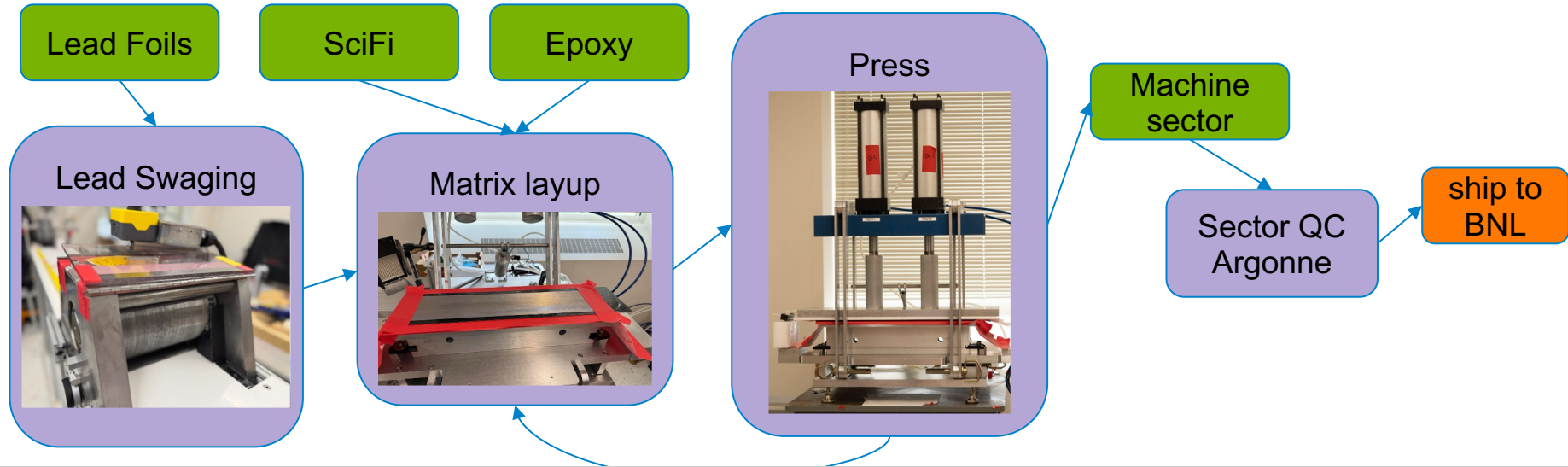


Tray Assembly	# parts	# per batch	# batches	hours/batch	minutes/pace	# people	Total Person hours	Inactive hours/batch	Batches/week/production line	# of production lines	Total production weeks	Total production years (w/ 85% annual efficiency)
assemble modules on tray	388	1	388	5	300	1	1940	0	8	3	16	0.37
Tray electrical QC	388	1	388	0.5	30	1	194	48	3	3	43	0.98
packing	388	1	388	0.5	30	2	388	0	80	3	2	0.04

# Workflow: Pb/SciFi Sectors & ESB



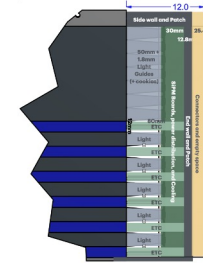
# Production Plan: Mechanical Sector



Sector Assembly	# parts	# per batch	# batches	hours/batch	minutes/piece	# people	Total Person hours	Inactive hours/batch	Batches/week/production line	# of production lines	Total production weeks	Total production years (w/ 85% efficiency)
Fiber QC (1%)	10000	208	48	52	15.0	2	104	104	0.8	1	62	1.4
Matrix layup	576	12	48	75	377	5	377	377	0.3	2	90	2.0
CF Frame Integration	288	6	48	16	160	2	32	80	1.3	2	19	0.4
Sector QC	48	1	48	8	480	2	16	16	5.0	1	10	0.2

# Production Plan: End-of-Sector Boxes

- component parts QC
- pre-assembly of ESB
  - SiPM's, CALOROC, cooling
- ESB QC
- Final assembly at BNL



Sector Assembly	# parts	# per batch	# batches	Active hours/batch	minutes/piece	# people	Total Person hours	Inactive hours/batch	Batches/week/production line	# of production lines	Total production weeks	Total production years (w/ 85% annual efficiency)
Light guide	5933	60	99	2	2	1	198	12	5	1	20	0.45
SiPM board QC	6022	60	101	2	2	1	202	12	5	1	22	0.5
CALOROC PCB QC	101	1	101	2	120	1	202	0	20	1	5	0.11
End-of-tray card QC	106	1	106	2	120	1	212	0	20	1	5.3	0.12
Cooling QC	100	1	100	0.5	30	1	50	0.5	40	1	2.5	0.06
ESB pre-assembly/QC	100	1	100	2	120	2	400	0	20	1	5	0.11